

B02

Progress of development of pnCCD image sensor onboard the next generation GRBs observation satellite mission HiZ-GUNDAM

Multi-messenger Astrophysics 2nd annual conference

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Division of Mathematical and Physical Sciences

Ryuji Kondo

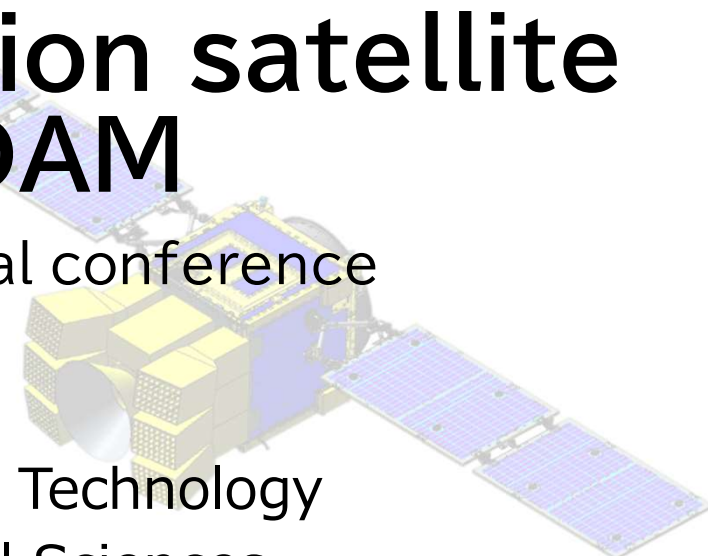
[S. Takahashi et al., NIM-A, 2024, Volume 1064, pages 169413](#)

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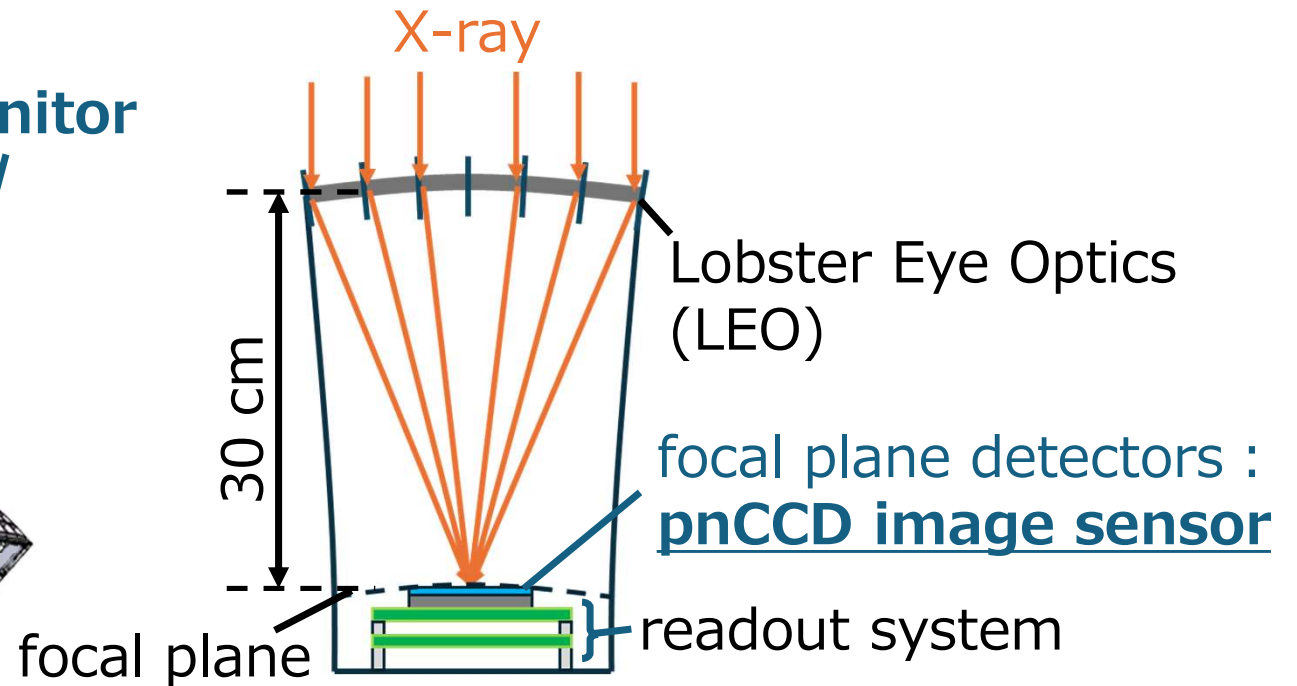
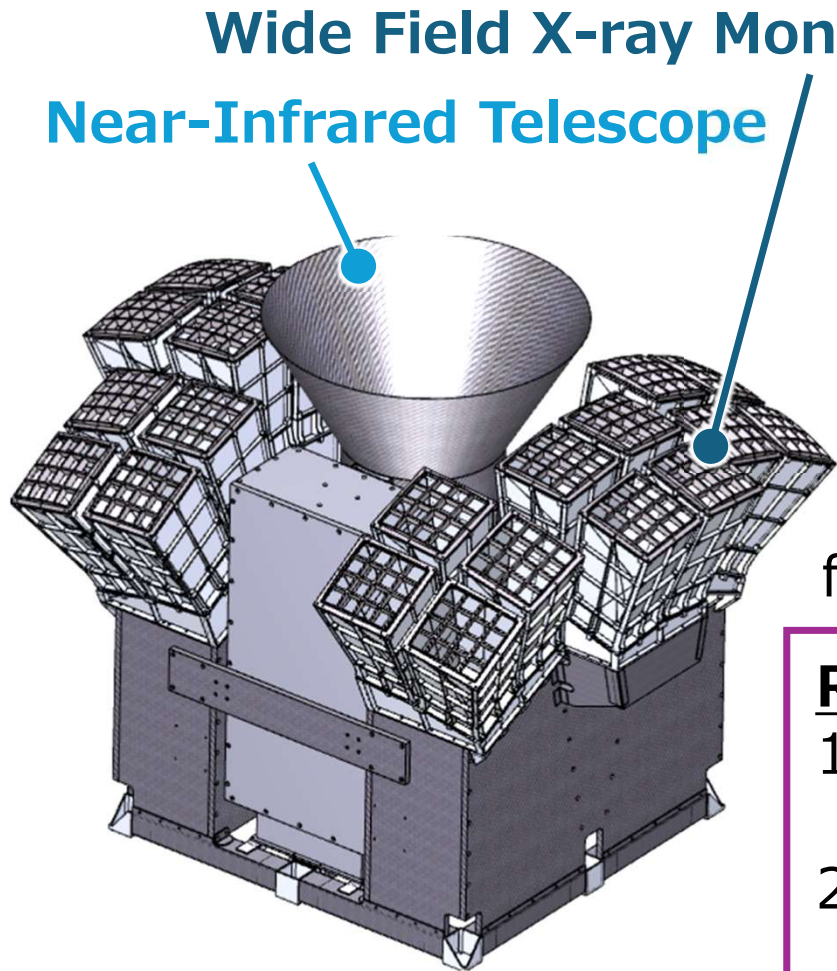
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Takanori Sakamoto⁴, Hsien-Chieh Shen⁴, Satoshi Hatori⁵, Ryoya Ishigami⁵, Akihiro Doi⁶, Hideo Matsuhara⁶

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HiZ-GUNDAM and Wide Field X-ray Monitor



Requirement for focal plane detectors

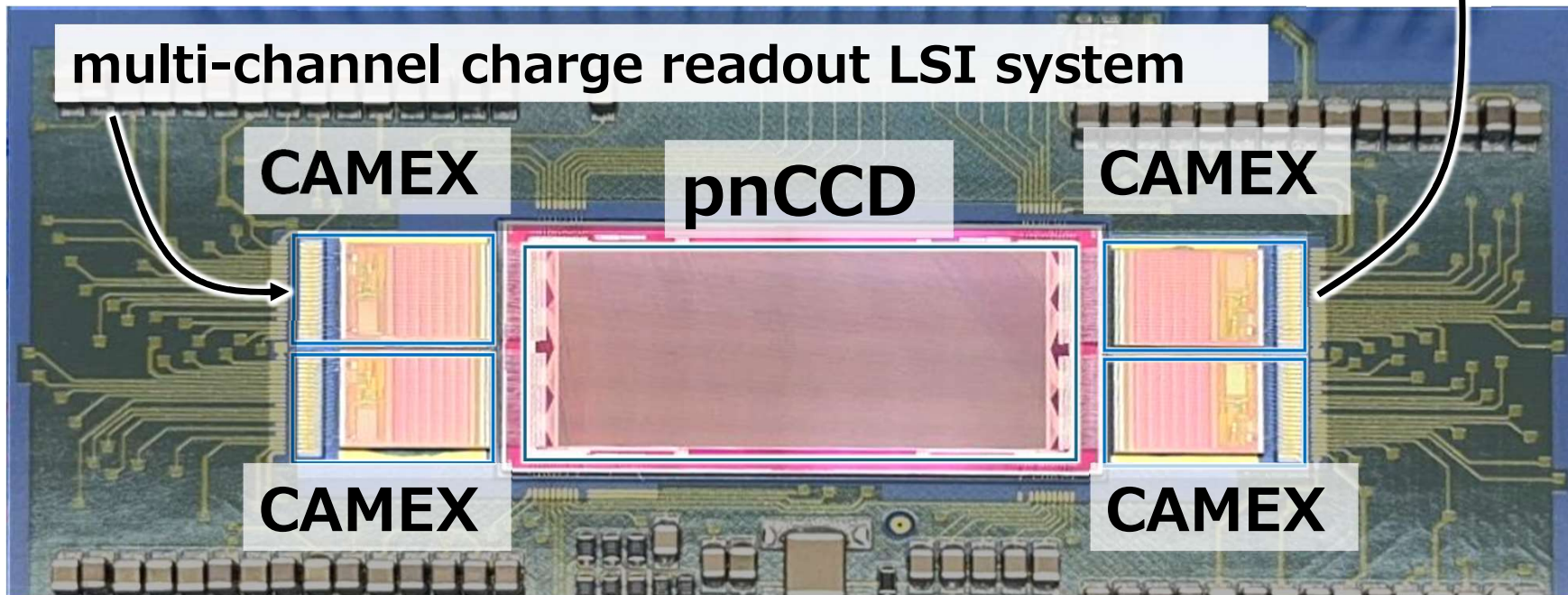
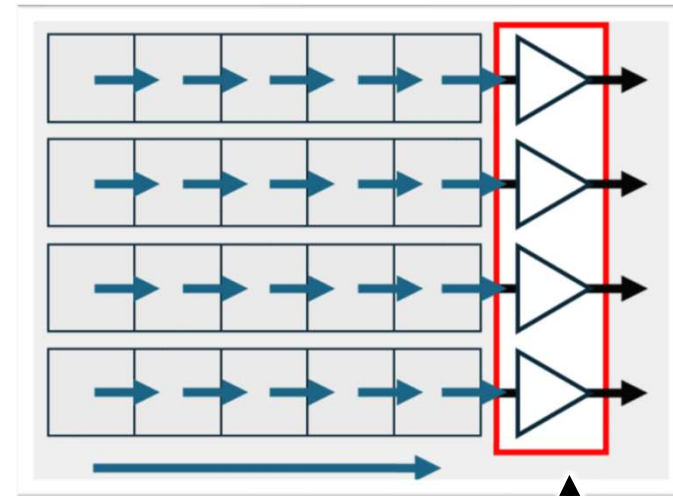
1. sensitivity range of LEO : 0.4 – 4 keV
Energy range covering **0.4 - 4 keV** sufficiently
2. Time resolution **shorter than 0.1 second** is required to observe time variability of GRBs

pnCCD Image Sensor

pnCCD image sensor

pnCCD system manufactured by
PNsensor GmbH

- **Multi-Channel Readout**
- **High frame rate**
(up to 1000 frames/second !)

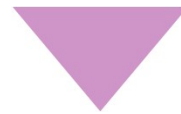


Achievements of pnCCD in Spacecrafts

SPOC / eROSITA (Germany)

SVOM/MXT (France)

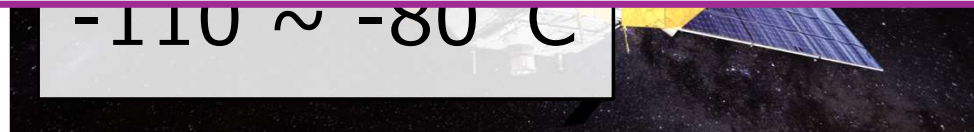
- *HiZ-GUNDAM* is M-class mission (~500 kg)
- Costly mechanical chillers cannot be installed.
- The system will be operated at about -30 °C with passive cooling only due to the strict power requirements.



Temperatures are higher than the thermal environment of pnCCDs in previous spacecraft.
→ **We need to investigate the effects of radiation on pnCCDs.**



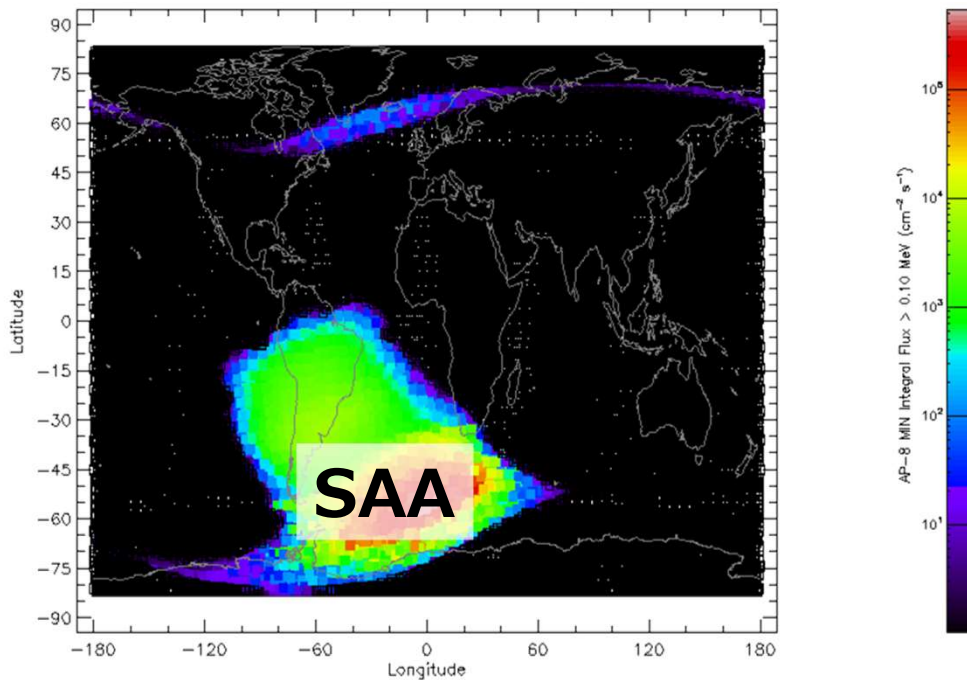
<http://sci.esa.int/xmm-newton/18015-xmm-newton-spacecraft/>



<https://ep.bao.ac.cn/ep/cms/files/W020210609570525377403.png>

Radiation and its effects on *HiZ-GUNDAM* in orbit

sun synchronous dawn-dusk orbit(altitude: 500 – 600 km, 98.1°)
mission term: 3 – 5 years



Mean Captured Proton Flux
(by SPENVIS)

Radiation Damage Effects on pnCCD

Increase in dark current
→ Worsen lower detectable energy



Requirement for
focal plane detector: **0.4** – 4 keV
may no longer be satisfied.

**We must test radiation tolerance
to various irradiation doses.**

Purpose

To quantify the change in soft X-ray detection performance of pnCCDs in response to changes in irradiation dose.

During the *HiZ-GUNDAM* mission, can the pnCCD operate with the **lower detectable energy=0.4 keV** requirement?

→ Investigate in the parameter space of **frame rate** and **temperature**.

pnCCD operating range
before irradiation

Items	Parameter
lower detectable energy	0.1 keV
frame rate	any
temperature	any

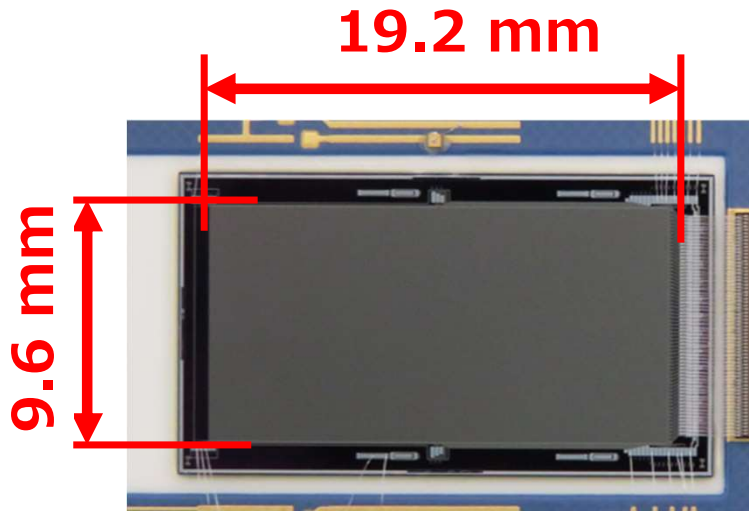


pnCCD operating range
after irradiation

Items	Parameter
lower detectable energy	≤ 0.4 keV
frame rate	\geq ??? frame/s
temperature	\leq ??? °C

Detector used in this study

- The focal plane detector of FM requires 55 mm × 55 mm sensitive area, but a small pnCCD image sensor was used for this study to test radiation tolerance.



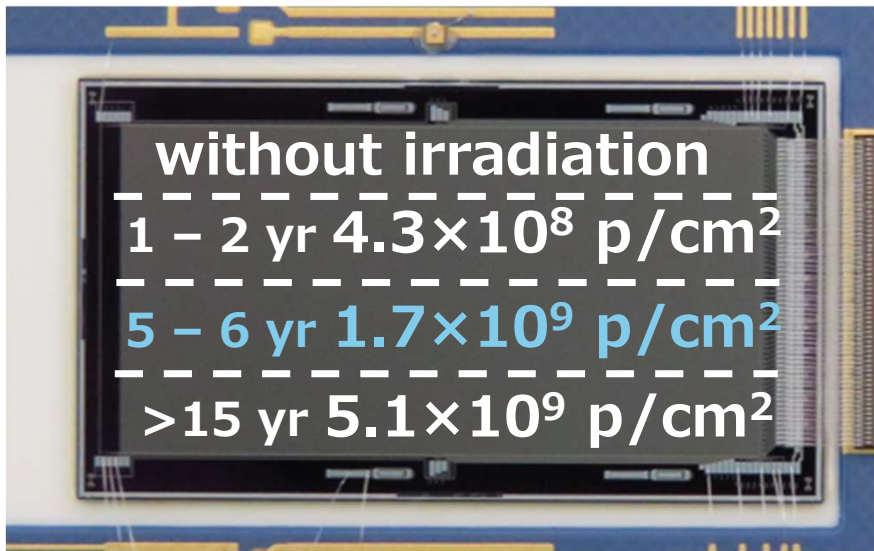
Item	small-pnCCD
sensitive area	9.6 × 19.2 mm ²
pixel size	75 μm
pixel number	128 × 256
depletion thickness	450 μm

Radiation Tolerance Test

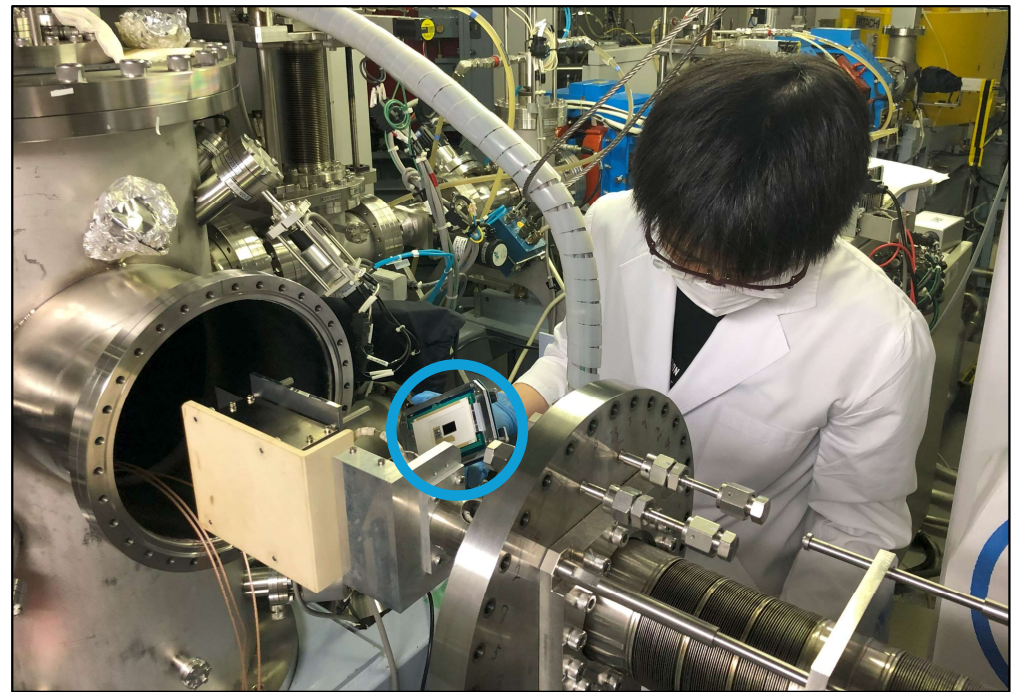
@ Wakasa-wan Energy Research Center (WERC)

The exposure received in orbit was represented by 10 MeV proton beams.

The pnCCD is divided into four irradiated areas, each of which is irradiated with a different number of protons.



pnCCD with separate irradiation area
(number of years in orbit in the **WORST** case
corresponding to the irradiation dose
→ Actual deterioration could be much less.)



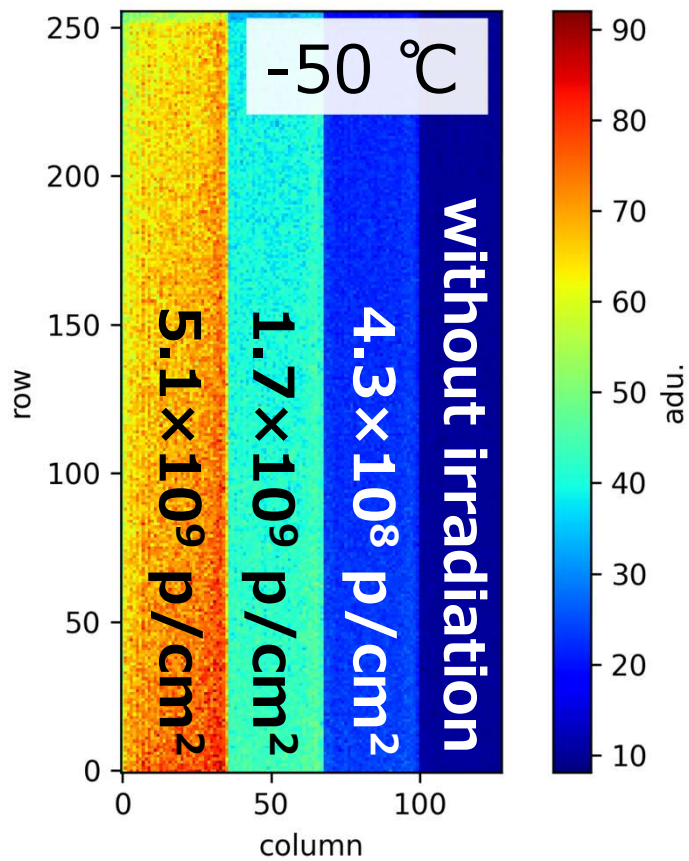
assembling the experimental setup 8/13

Image data after irradiation

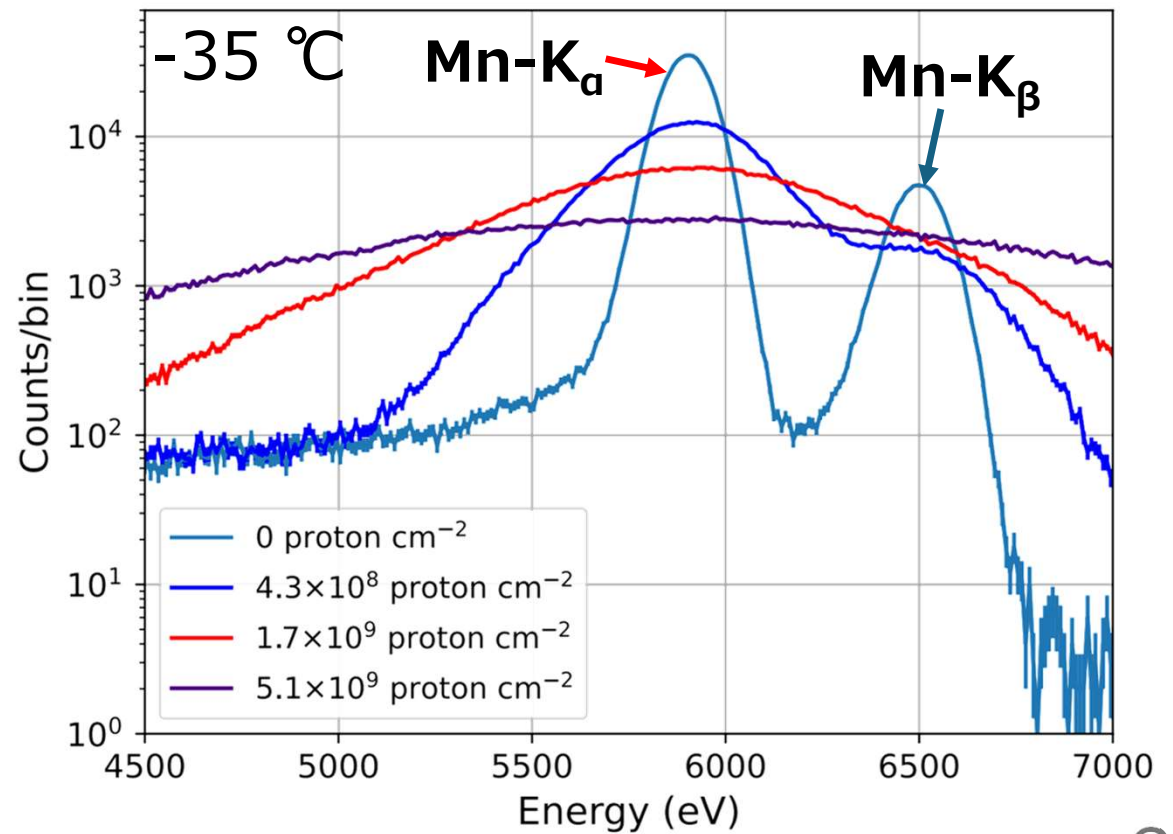
frame rate: 33 frames/second

temperature: -20°C , -35°C , -50°C , -70°C , -90°C

Dark noise (e^-)

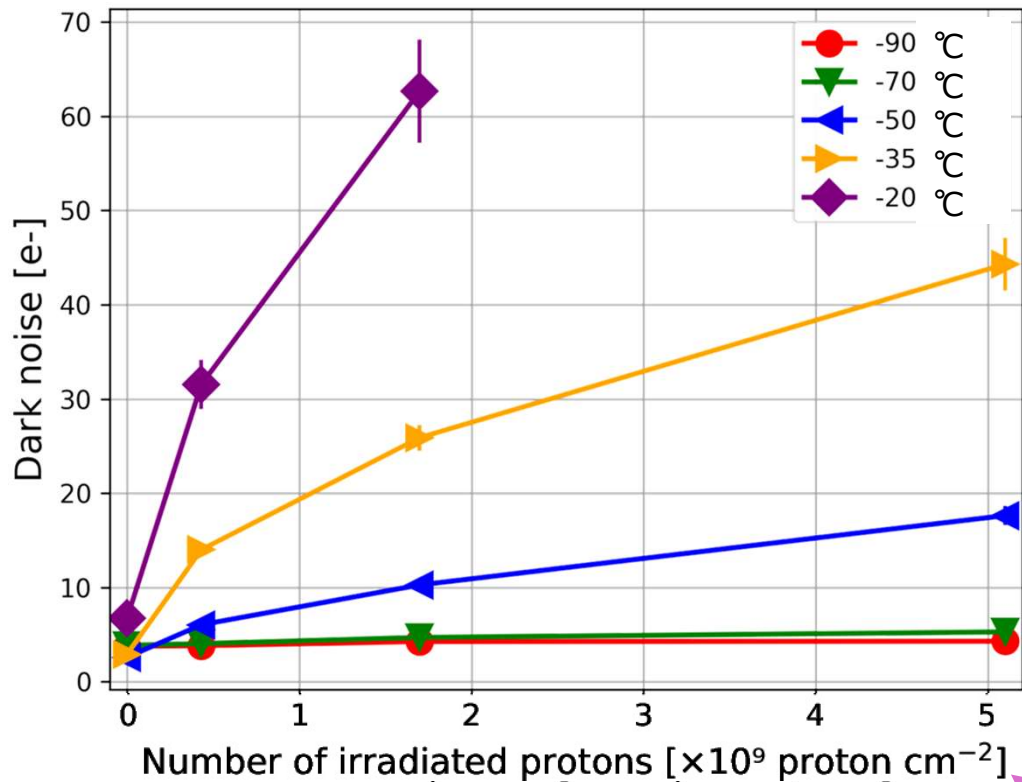


^{55}Fe (Mn- K_{α} :5.9 keV, Mn- K_{β} :6.5 keV)

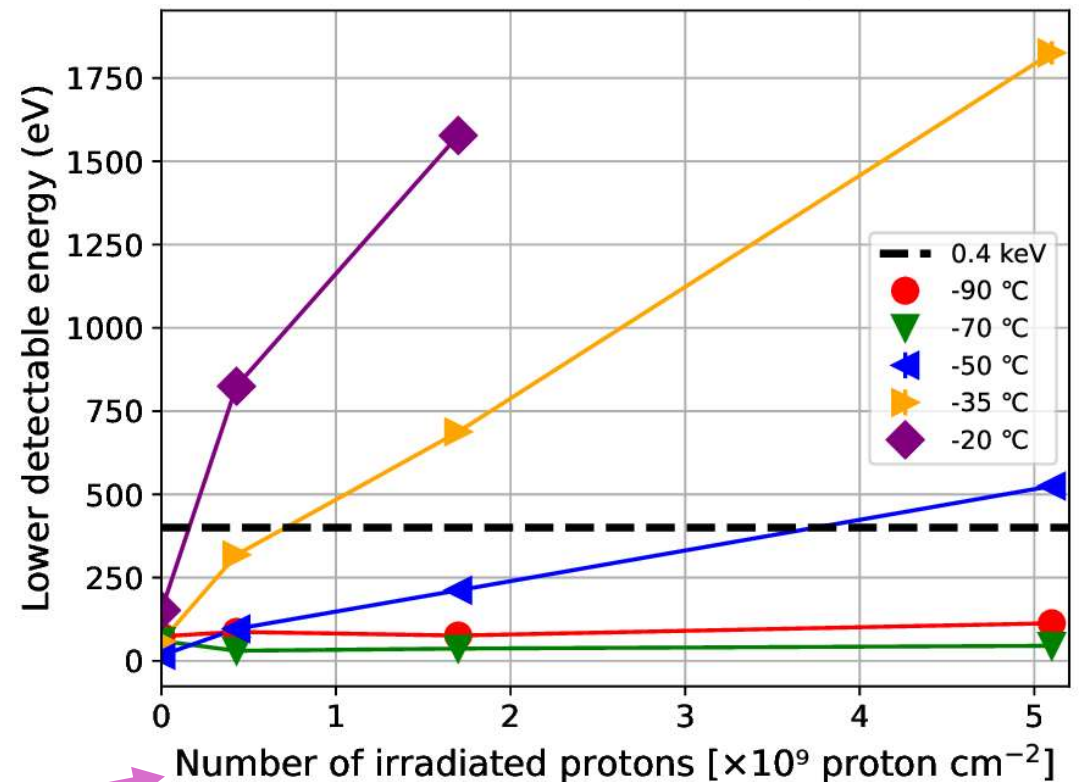


Dark Noise and Lower Detectable Energy

Dark Noise



Lower Detectable Energy

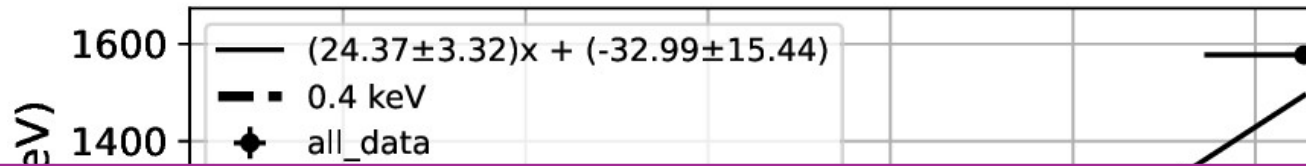


Frame rate: 33 Hz

identical

A correlation can be plotted between dark noise and lower detectable energy.

Dark noise v.s. Lower Detectable Energy

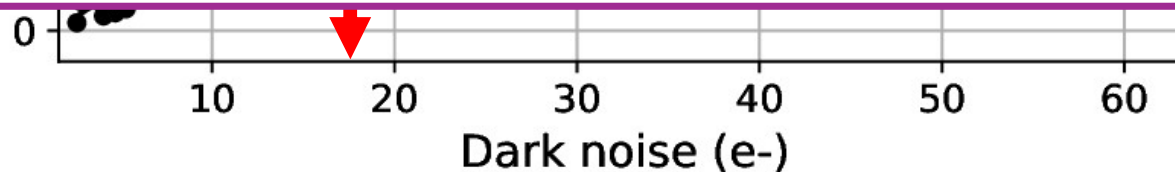


Assume the target value for dark noise is **18 e-**,

$$(\text{Dark noise}) \propto t_{\text{exposure}}$$

$$\propto (\text{frame rate})^{-1}$$

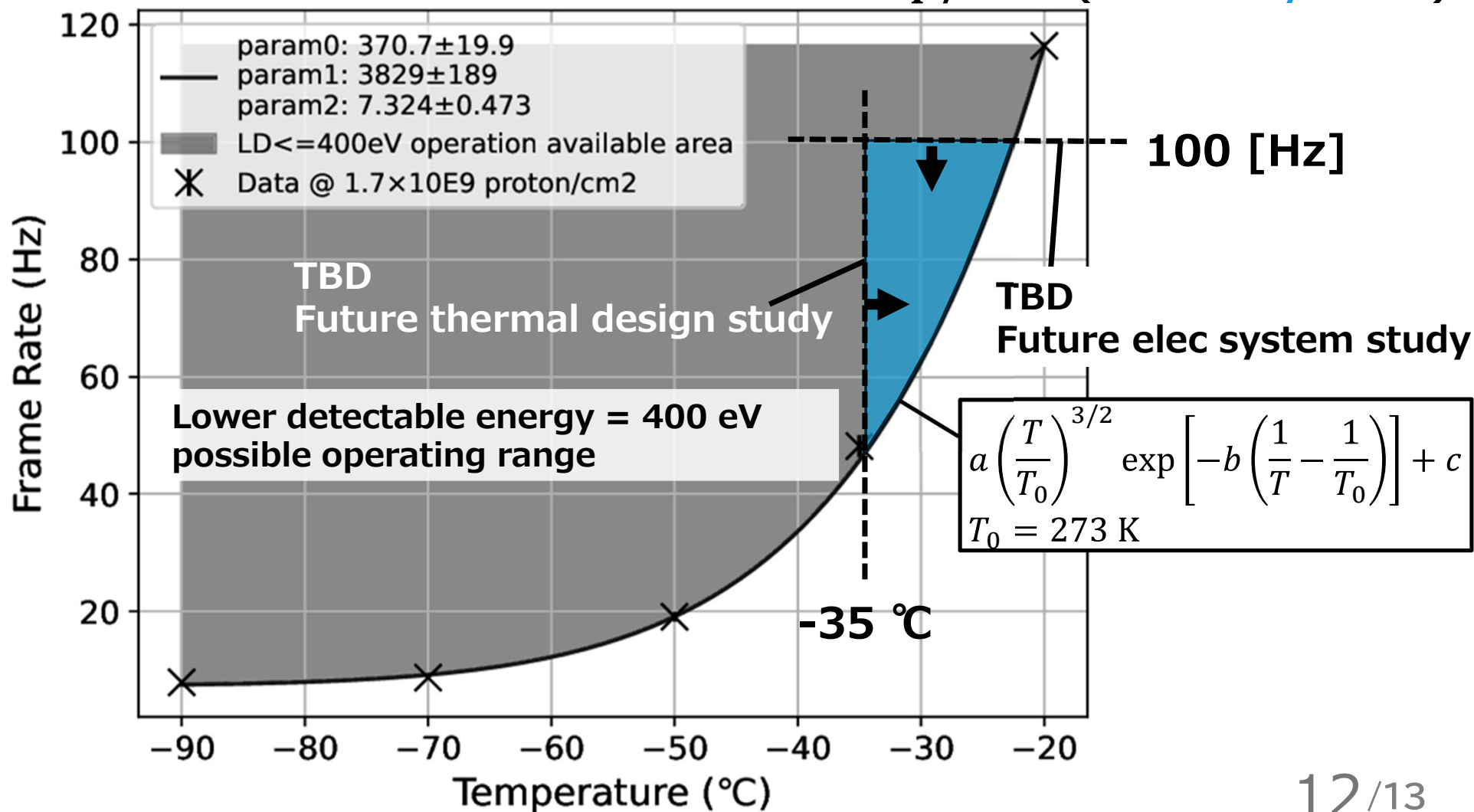
Using this relationship, we searched for an acceptable parameter space for temperature and frame rate.



Dark noise, where lower detectable energy is 400 eV of the required value, is **~18 e-**.

Temperature v.s. Frame Rate

@ 1.7×10^9 p/cm² (5 – 6 years)



Summary and Future

Summary

- In this study, radiation tolerance tests were conducted on pnCCDs equipped with *HiZ-GUNDAM*.
- For four different exposures and arbitrary temperatures (-90°C to -20°C), the range of frame rates over which a pnCCD can operate with lower detectable energy = 0.4 keV requirement was investigated.
- The results show that *HiZ-GUNDAM* can operate with lower detectable energy = 0.4 keV for the mission duration.

Future

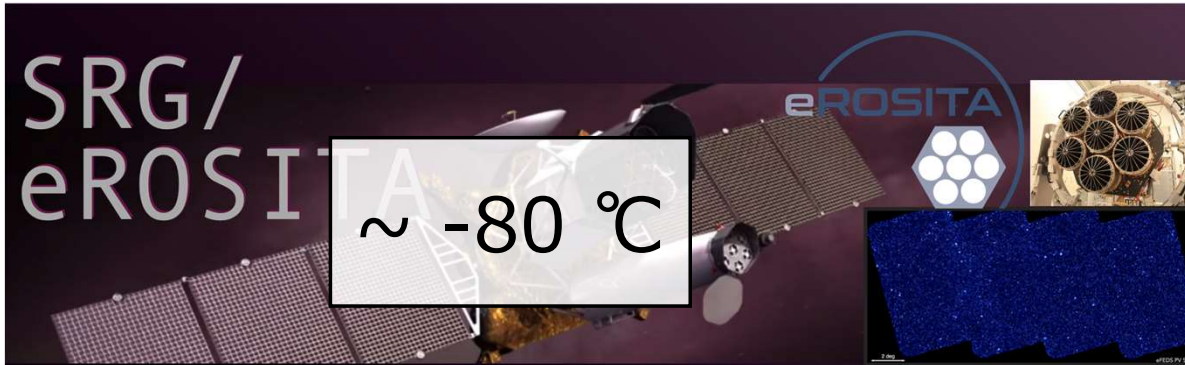
Starting with thermal design and electrical system considerations, operating temperatures and frame rates will be discussed.

[S. Takahashi et al., NIM-A, 2024, Volume 1064, pages 169413](#)

APPENDIX

Achievements of pnCCD in Spacecrafts

SRG/eROSITA(Germany)



~ -80 °C

https://heasarc.gsfc.nasa.gov/docs/srg/erosita/inc/erosita_banner.png

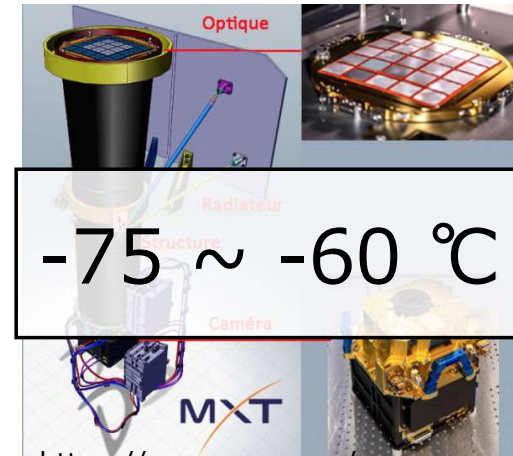
XMM-Newton(Europe)



~ -90 °C

<http://sci.esa.int/xmm-newton/18015-xmm-newton-spacecraft/>

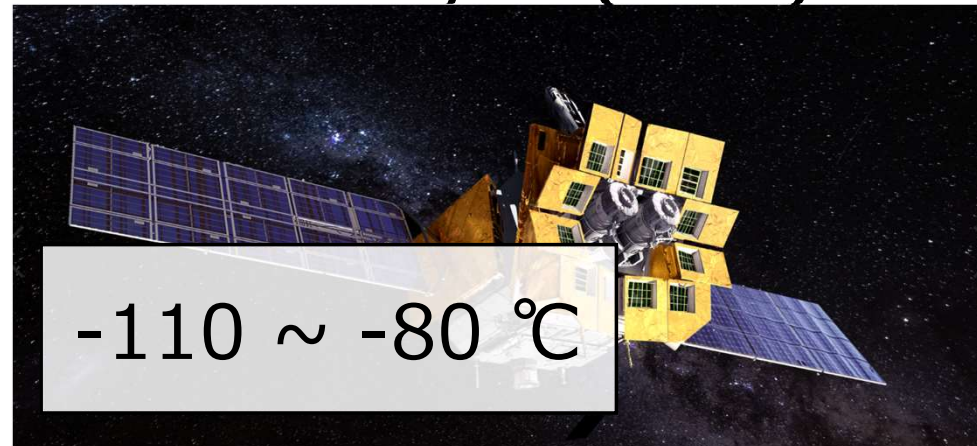
SVOM/MXT(France)



-75 ~ -60 °C

https://www.svom.eu/wp-content/uploads/2021/06/4922_5-1024x926.jpg

Einstein Probe/FXT(China)

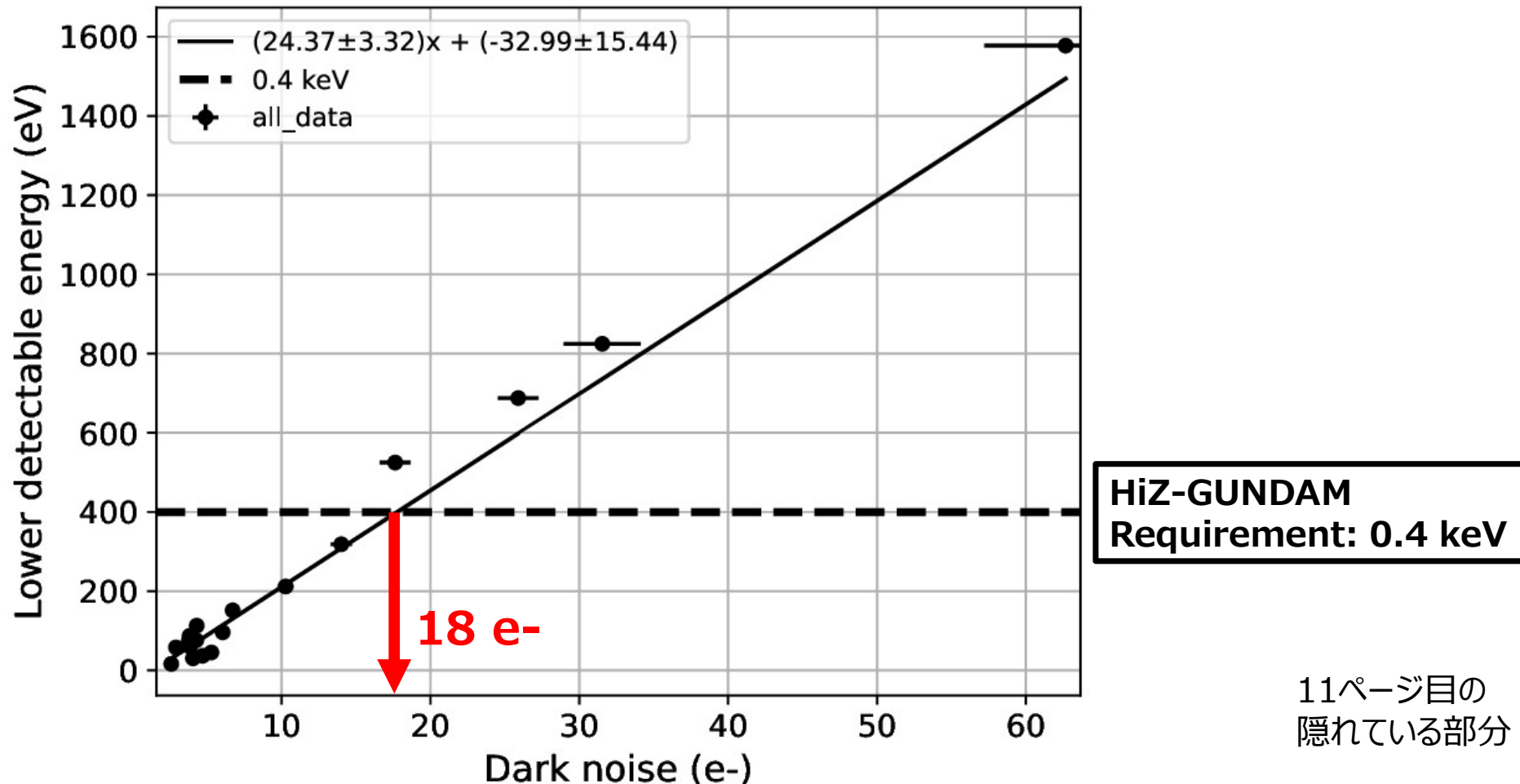


-110 ~ -80 °C

<https://ep.bao.ac.cn/ep/cms/files/W020210609570525377403.png>

4ページ目の
隠れている部分

Dark noise v.s. Lower Detectable Energy



Dark noise, where lower detectable energy is 400 eV of the required value, is **~18 e-**.

5 σ にした理由

X線バックグラウンドレート

- HiZ-GUNDAMでは宇宙X線背景放射によるバックグラウンドレートは露光時間を0.1sで

0.355 counts/frame

ノイズ誤検知

- 5 σ の有意度で信号を検出する場合、ノイズ成分を信号として誤検出する確率が、

$$3 \times 10^{-5} \text{ event/s} \times 700^2 \text{ pixels/frame} \times 0.1 \text{ s} \simeq \mathbf{0.1 \text{ event/frame}}$$

ノイズの誤検知率がバックグラウンドレートと同等以下となるため5 σ を採用している。

モデル関数の導出

33 Hzでのダークノイズとフレームレートの関係

$$(\text{Dark noise_@33 Hz}) \propto (\text{Frame rate} = 33 \text{ Hz})^{-1}$$

400 eVを満足するダークノイズとフレームレートの関係

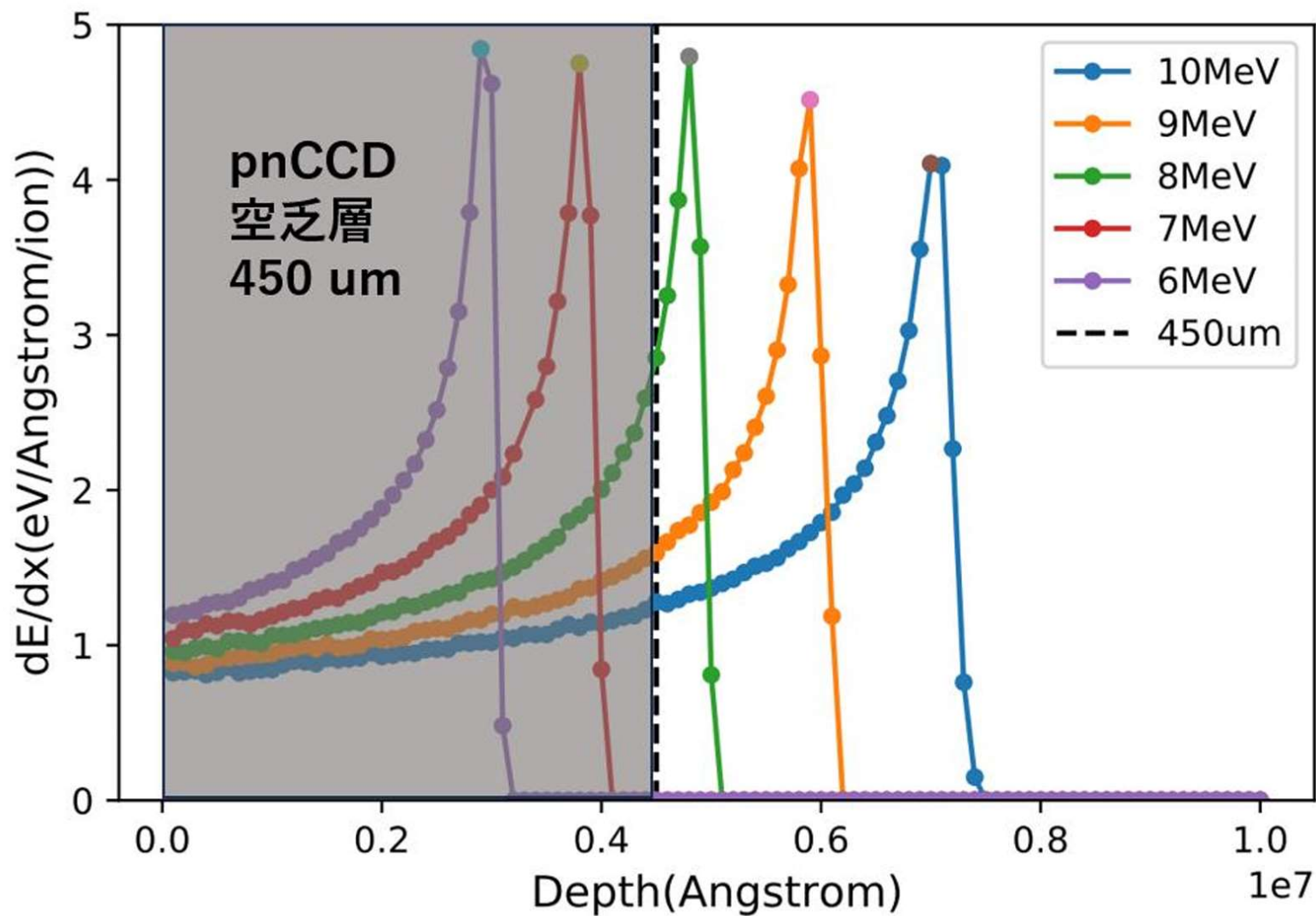
$$(\text{Dark noise_@400 eV}) \propto (\text{Frame rate_@400 eV})^{-1}$$

$$\frac{(\text{Dark noise_@33 Hz})}{18 e^-} \propto \left(\frac{33 \text{ Hz}}{(\text{Frame rate_@400 eV})} \right)^{-1}$$

最小検出エネルギー=400 eV に対応する目標値

$$\begin{aligned} (\text{Frame rate_@400 eV}) &\propto (\text{Dark noise_@33 Hz}) \\ &\propto T^{\frac{3}{2}} \exp\left(-\frac{Eg}{2k_B T}\right) \end{aligned}$$

10 MeV陽子線を選んだ理由

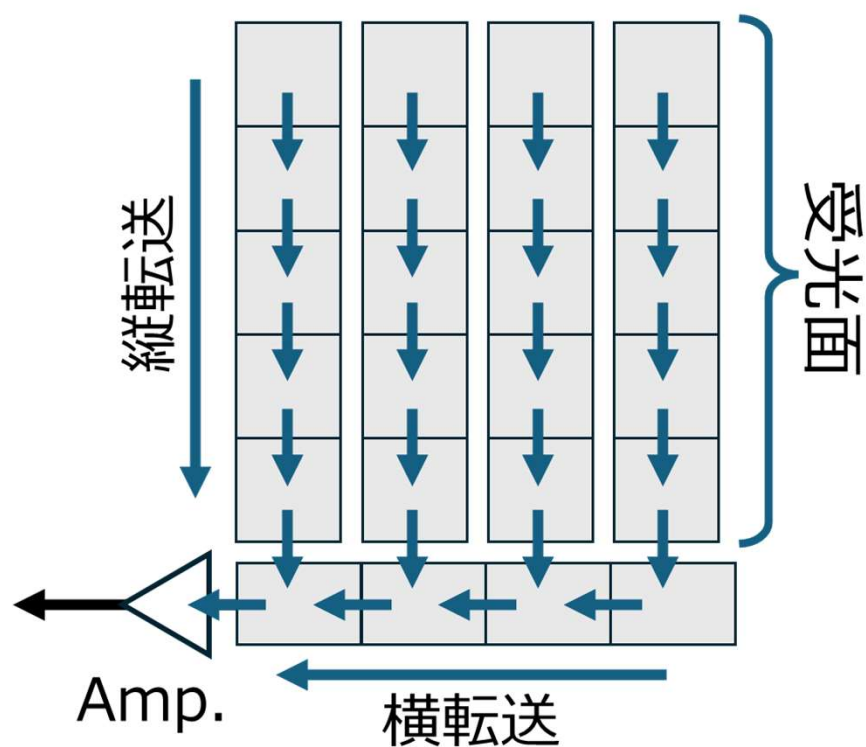


一般的なCCDとpnCCDの違い

一般的なCCD

1読み出しチャンネル

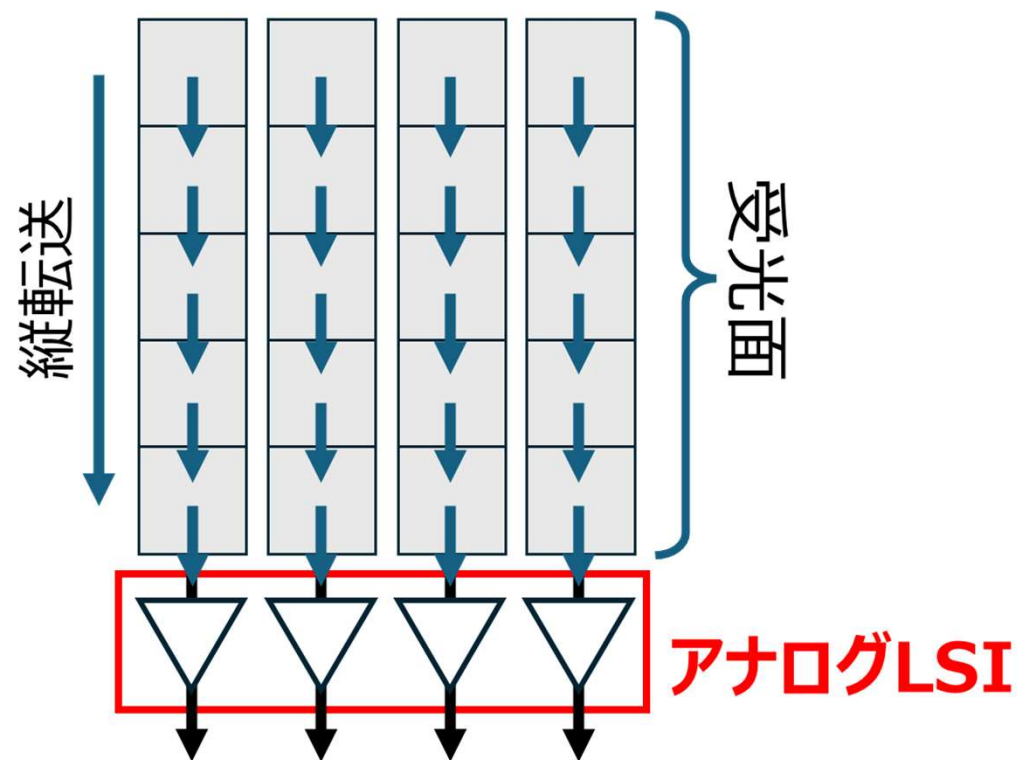
→ 遅い (~0.1-1 fps)



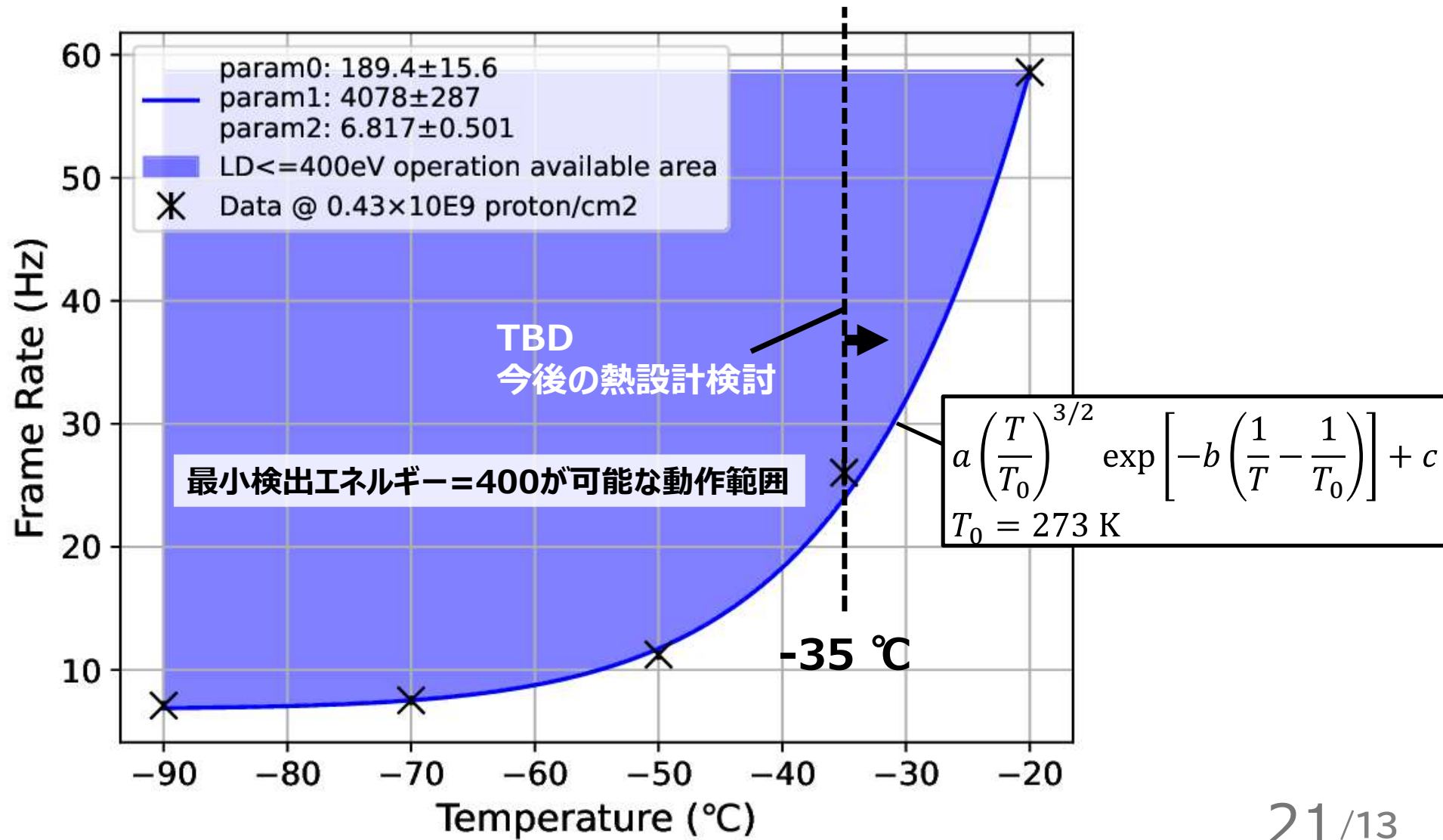
pnCCD

全列に読み出しチャンネル

→ **速い!!! (~10-1000 fps)**



温度 v.s. フレームレート@ 4.3×10^8 p/cm²(1-2年分)



Temperature v.s. Frame Rate

@ 1.7×10^9 p/cm² (5 – 6 years)

