

Simulating Earth-limb Gamma Rays

Ben Y. O. Xu ¹ Kenny C. Y. Ng ¹ Chingam Fong ¹ Zhe Li ²

¹Department of Physics, The Chinese University of Hong Kong

²Institute of High Energy Physics, Chinese Academy of Sciences

benxu@link.cuhk.edu.hk

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Introduction

- Earth Atmospheric γ -rays
- Motivation

Simulation Setup

- Back-tracing
- Interaction Simulation

Results

- Energy Spectrum
- Angular Distribution

Appendix

- Production Mechanism
- Geographical Distribution of γ -intensity

Earth Atmospheric γ -rays

- Produced by interactions between cosmic rays (CR) and the Earth's atmosphere.

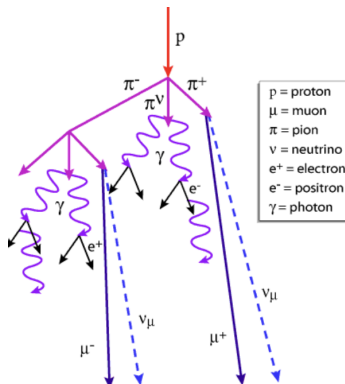
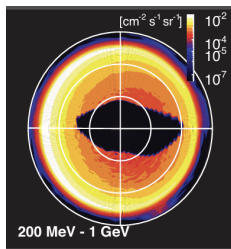


Image from <https://www.hawc-observatory.org/science/cosmicrays.php>

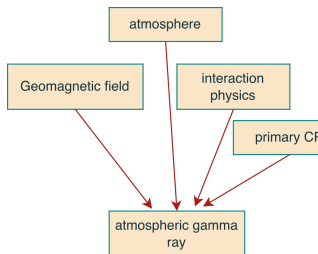
Motivation for Simulating Atmospheric γ -rays

- Atmospheric γ -rays act as a strong background for astronomical observations

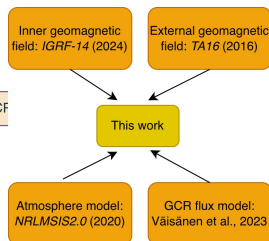


γ -ray intensity viewed from Fermi-LAT (Abdo et al., 2009)

- Testing ground for physics models

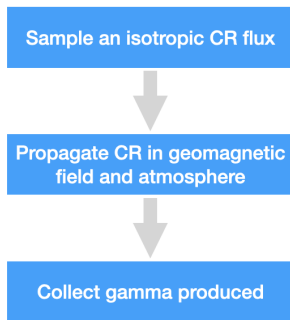


- No simulations using modern models have been conducted in over two decades

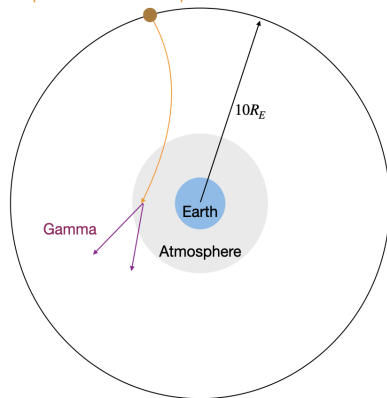


A Simple Simulation

- ▶ A naive approach to the simulation:



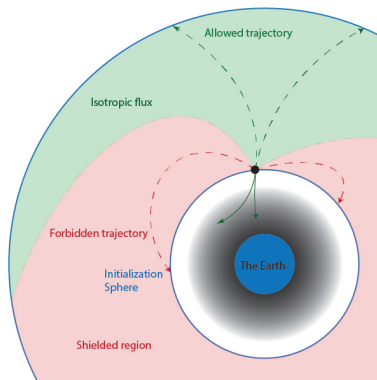
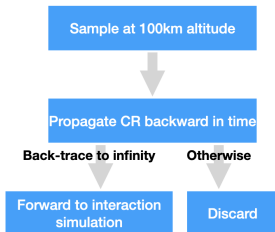
Isotropic distribution of CR primaries



- ▶ However, many cosmic rays miss the atmosphere and do not produce γ -rays.

Back-tracing

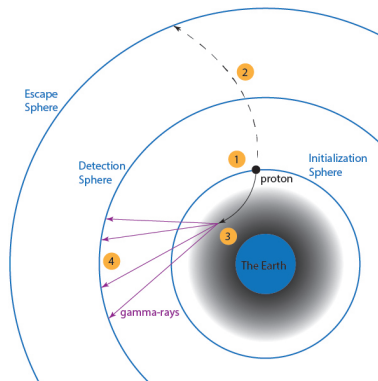
- ▶ Geomagnetic field modeled up to a distance of $10R_E$ (Escape sphere) using:
 - ▶ Internal geomagnetic field: IGRF-14 model (International Association of Geomagnetism and Aeronomy, 2024)
 - ▶ External field: TA16 (Tsyganenko et al., 2016)
- ▶ Liouville's theorem: if the particle flux at infinity is isotropic, the particle flux at an arbitrary point is the same as at infinity along a set of allowed trajectories, and 0 along the forbidden trajectories
- ▶ Recipe:



A simplified illustration

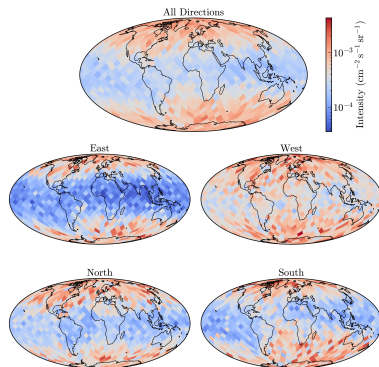
Interaction Simulation

- ▶ The atmosphere is modeled as 200 concentric spherical shells extending from sea level to an altitude of 100 km
- ▶ Each shell has a uniform thickness of 0.5 km
- ▶ Shells are divided into an 18×18 grid with equal intervals of latitude (10°) and longitude (20°)
- ▶ The NRLMSIS2.0 atmospheric model (Emmert et al., 2021) calculates the densities of hydrogen, helium, oxygen, nitrogen, and argon at each grid's midpoint
- ▶ Particle interactions are modeled using the FTFP_BERT physics list in GEANT4 (Allison et al., 2016)



Benchmarking Against Fermi-LAT Observations

- ▶ Fermi LAT γ -ray observations: between June and September 2008 (Abdo et al., 2009)
- ▶ Simulation parameters: using input parameters corresponding to 00 : 00 UTC on 12th June 2008
- ▶ Simulated γ -ray flux was extracted at Fermi's orbital altitude and inclination:
 - ▶ Altitude = 565 km (Atwood et al., 2009)
 - ▶ Inclination within 25.5° (Atwood et al., 2009)

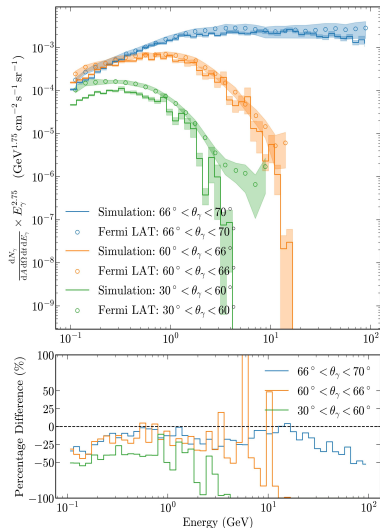
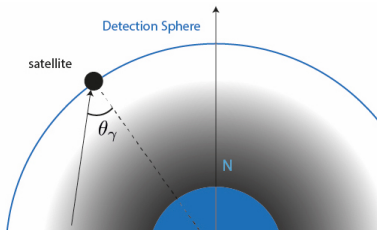


Remarks:

- ▶ Fermi-LAT does not record data inside the South Atlantic Anomaly (SAA) (Atwood et al., 2009).
- ▶ Orbital precession introduces non-uniform exposure across geographic locations and pointing directions.
- ▶ Our current simulation neither masks the SAA region nor applies exposure weighting to the flux.

Energy Spectrum

- ▶ Reproduces the general spectral shape across all evaluated nadir angle (θ_γ) intervals
- ▶ $\sim 25\%$ discrepancy at $66^\circ < \theta_\gamma < 70^\circ$ and $60^\circ < \theta_\gamma < 66^\circ$
- ▶ $\sim 50\%$ discrepancy at $30^\circ < \theta_\gamma < 60^\circ$
- ▶ The exact cause of this systematic discrepancy remains under investigation



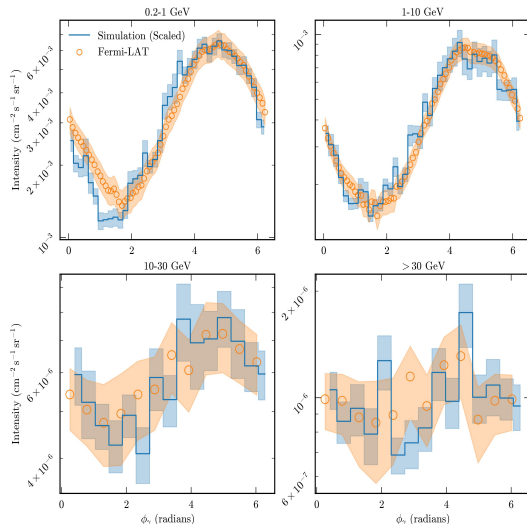
East-West Effect

North	$0^\circ < \phi_\gamma < 45^\circ$ $315^\circ < \phi_\gamma < 360^\circ$
East	$45^\circ < \phi_\gamma < 135^\circ$
South	$135^\circ < \phi_\gamma < 225^\circ$
West	$225^\circ < \phi_\gamma < 315^\circ$

- γ -rays selected from the $60^\circ < \theta_\gamma < 75^\circ$ range
- Simulated intensities are scaled:

0.2-1 GeV	1-10 GeV	10-30 GeV	>30 GeV
1.19	1.23	1.14	1.58

- General agreement in shape with observed data

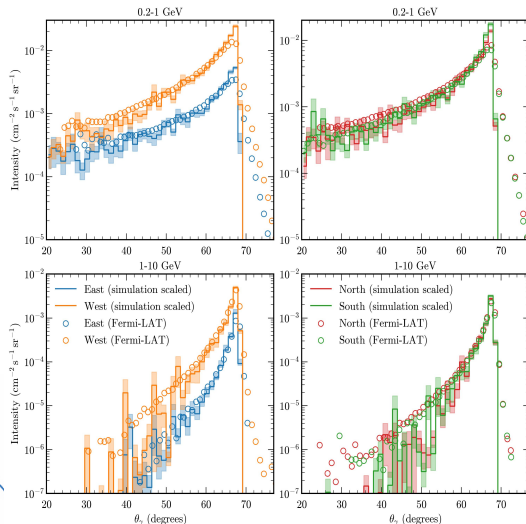
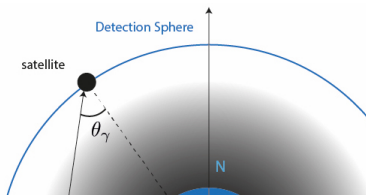


Nadir Angle Distribution

- Simulated intensities are scaled:

0.2-1 GeV	1-10 GeV
1.35	1.14

- General agreement in shape
- The observed distribution is slightly smeared due to the Fermi-LAT's point spread function



Conclusion

- ▶ Developed a simulation for Earth's atmospheric γ -rays.
- ▶ The simulation successfully reproduces energy spectra and angular distributions that match observations fairly well
- ▶ A discrepancy of 25 – 50% remains in the overall normalization
- ▶ Future Work:
 1. Use the simulation to investigate the relationship between γ -ray intensity and proton cutoff rigidity, as observed in Madlee et al., 2020.
 2. Extend the study to other secondary particles (e.g., neutrinos, neutrons).

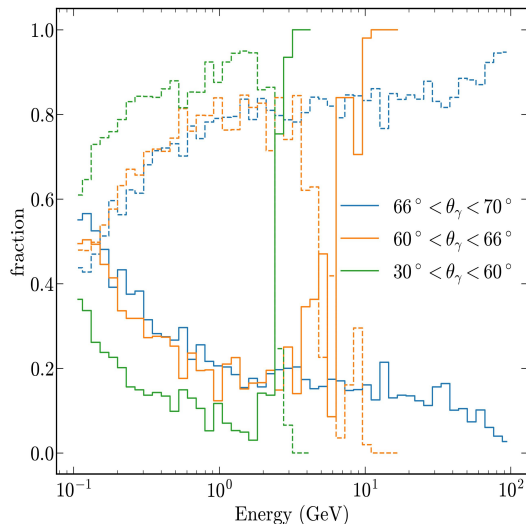
Acknowledgement

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Production Mechanism

- **Dashed lines:** Decay processes
- **Solid lines:** Bremsstrahlung radiation



Geographical Distribution of γ -intensity

- ▶ γ -rays selected at:
 - ▶ 565 km altitude
 - ▶ $60^\circ < \theta_\gamma < 75^\circ$

North	$0^\circ < \phi_\gamma < 45^\circ$ $315^\circ < \phi_\gamma < 360^\circ$
East	$45^\circ < \phi_\gamma < 135^\circ$
South	$135^\circ < \phi_\gamma < 225^\circ$
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