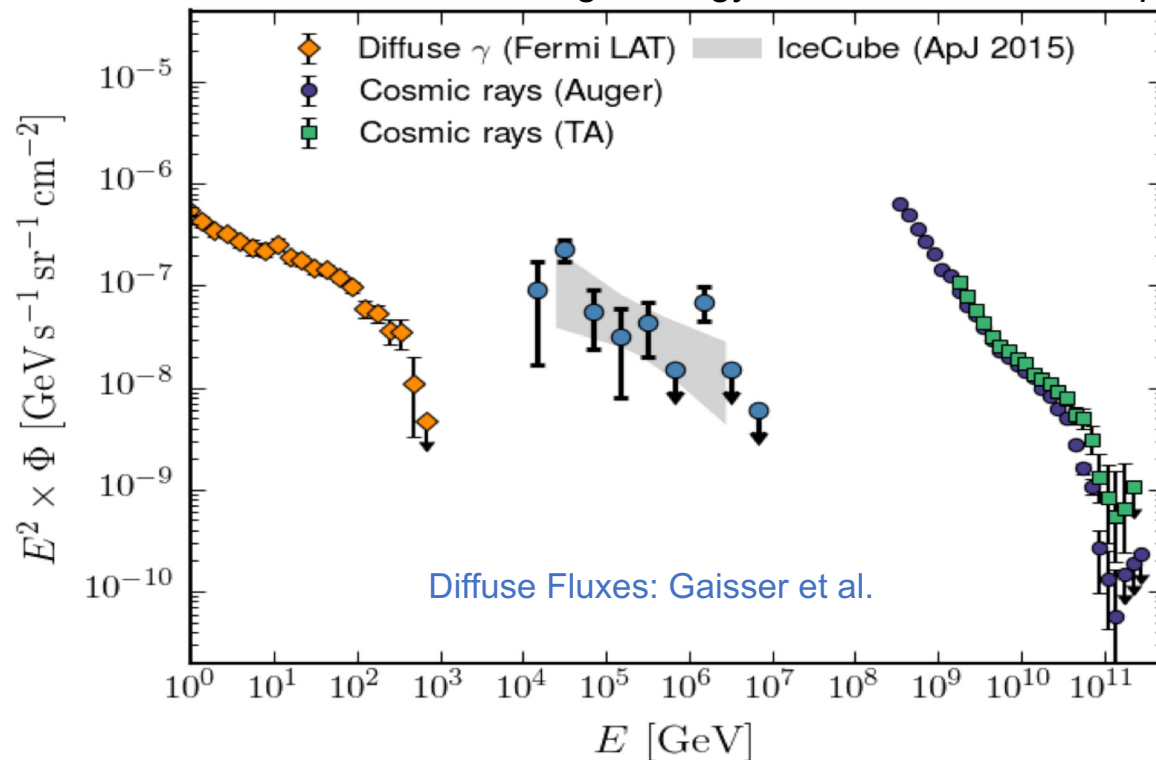


Elisabete Maria de Gouveia Dal Pino (IAG – USP, Brazil)



Motivation

High-energy diffuse fluxes are comparable



MAYBE

Galaxy Clusters hosting AGN
are suitable candidates for diffuse
multi-messenger:
size ~ 1 Mpc, $B \sim 10^{-6}$ G,
 $n \sim 10^{-2} \text{ cm}^{-3}$ & $T \sim 10^8$ K

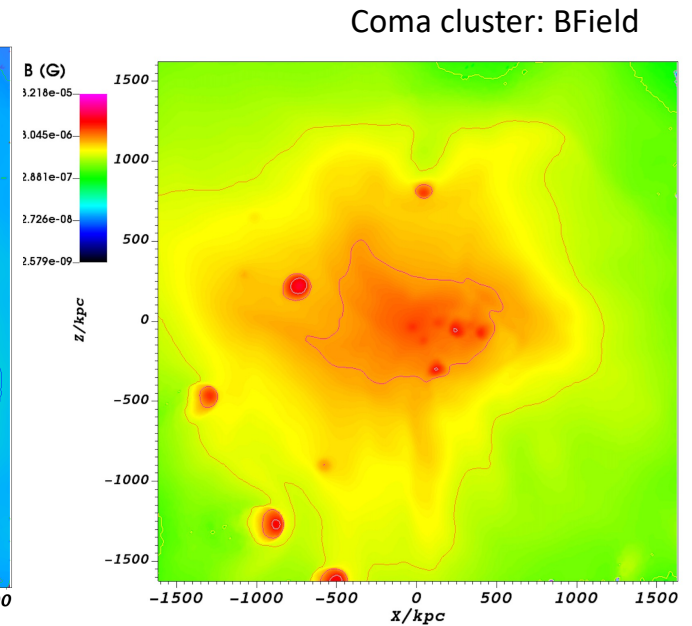
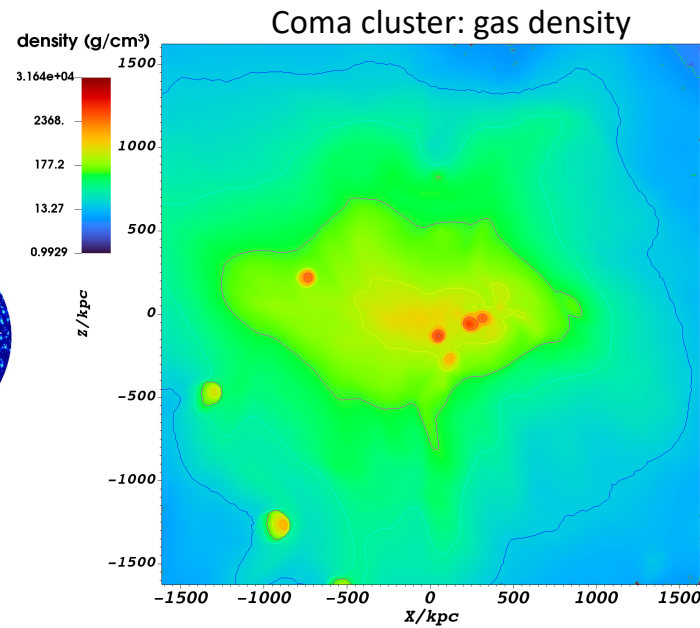
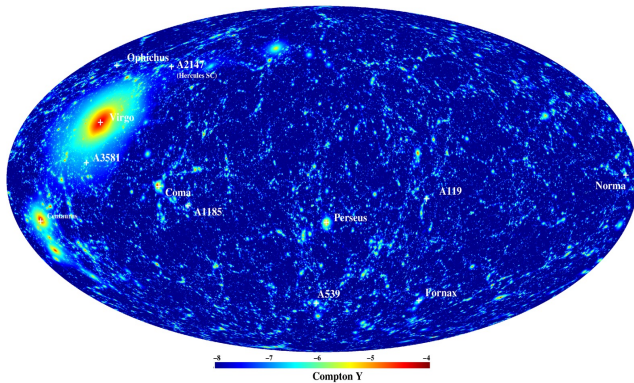
- ❖ *CRs, neutrinos and gamma rays have common origin?*
- ❖ *Single class of astrophysical sources can produce them?*

MHD Simulations of Galaxy Clusters hosting AGN

MHD simulation provides the background environment of galaxy clusters across redshift

**MHD SLOW simulation: cross identified
local galaxy clusters within 500 Mpc**

Dolag et al. 2023



Saqib Hussain (saqib.hussain@ung.si)

MHD + Monte Carlo simulations

SLOW Sim.: Magnetic-field, Gas density, Bremsstrahlung radiation

Background photo fields: CMB, EBL, and Radio background

Monte-Carlo code: CRPropa3

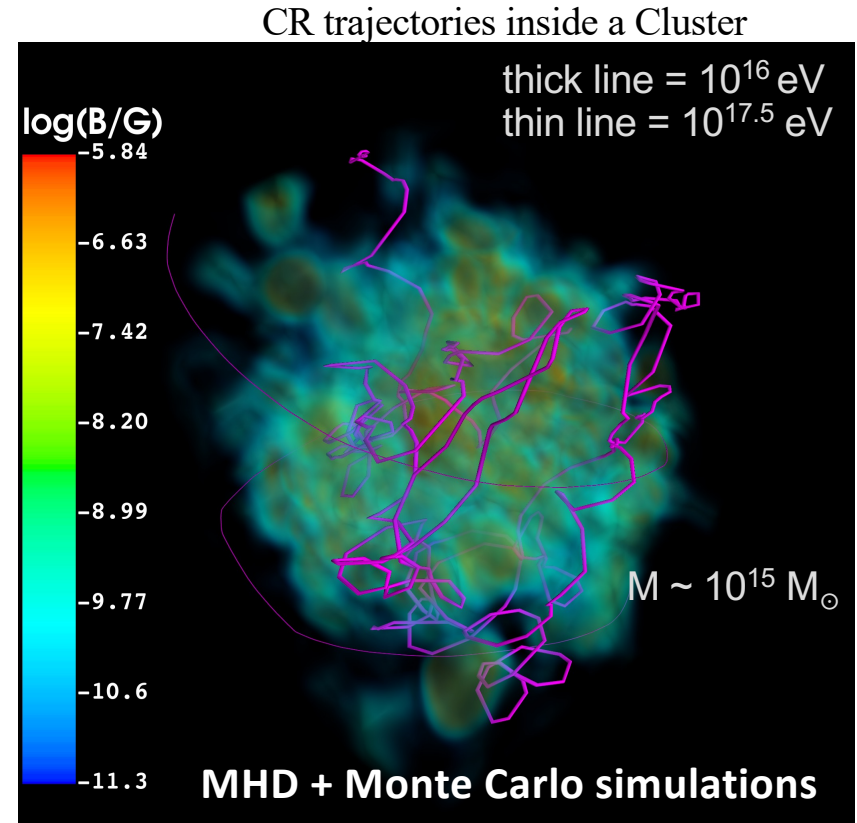
Cosmic Ray Propagation
(Hadronic and leptonic)

Interactions: photohadronic,
photonuclear, hadronuclear, and
electromagnetic cascade

Low energy ($< 10^{17}$ eV) CRs: Diffusive

High Energy ($> 10^{17}$ eV) CRs:
Semi-diffusive or Ballistic

Computationally very expensive: Tackled with AI/ML



Hussain et al. Nat. Commun.

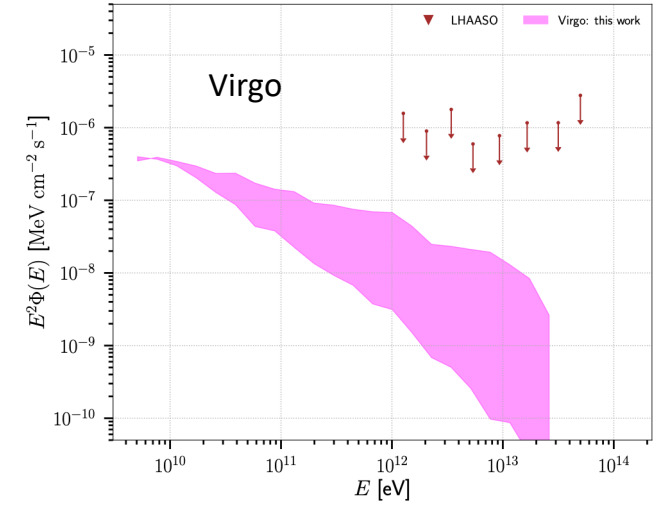
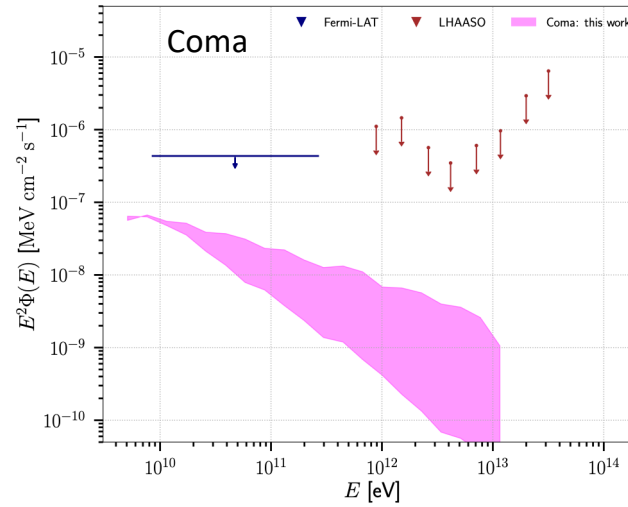
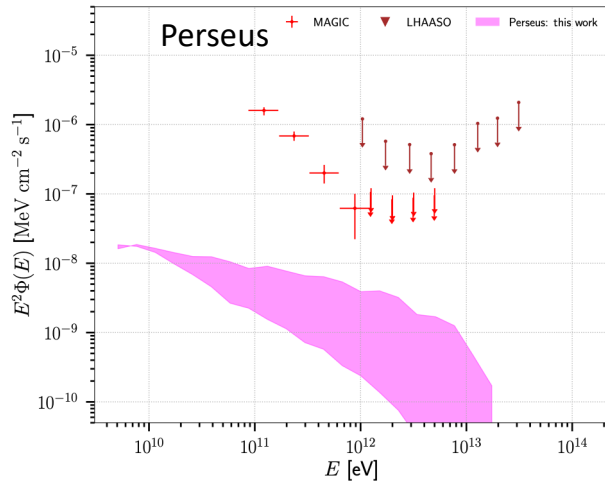
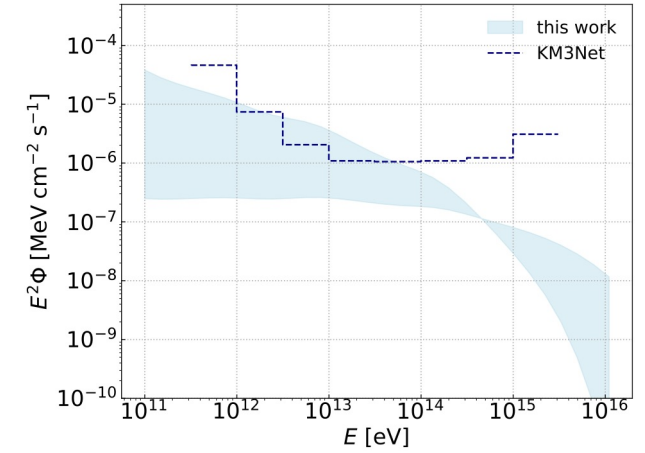
Gamm rays and neutrino from individual clusters

Parameters:

Index = 2.0 - 2.5

$E_{\text{max}} = 10^{16} - 10^{17}$ eV

$X_{\text{CR}} = E_{\text{CR}} / E_{\text{thermal}} = 0.01$

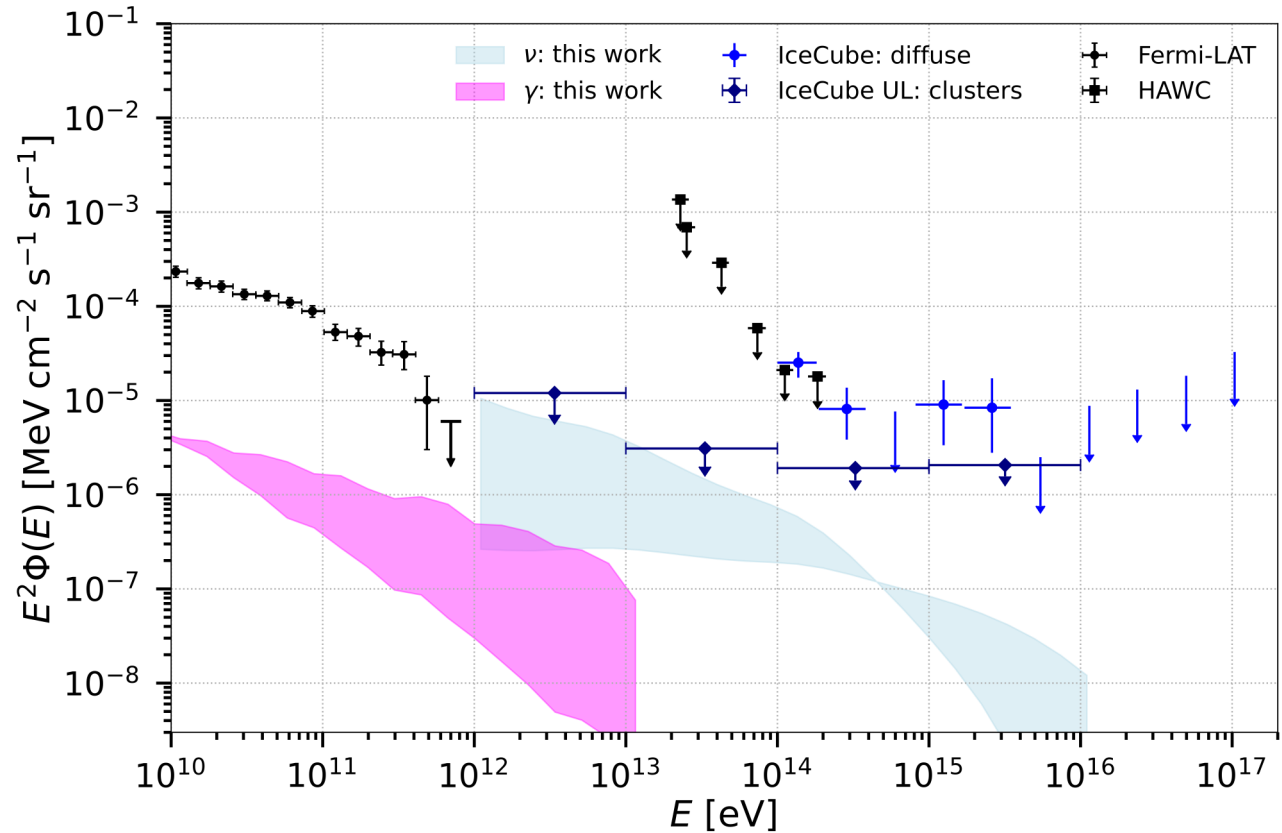


Integrated Gamma-ray and Neutrino Emission from Galaxy Clusters

Galaxy Clusters (GC):
within 500 Mpc & $M > 10^{14} M_{\text{sun}}$

GC contribution to diffuse
Gamma-ray background is limited

GC contribution to diffuse
Neutrino background is comparable
with the IceCube upper limits

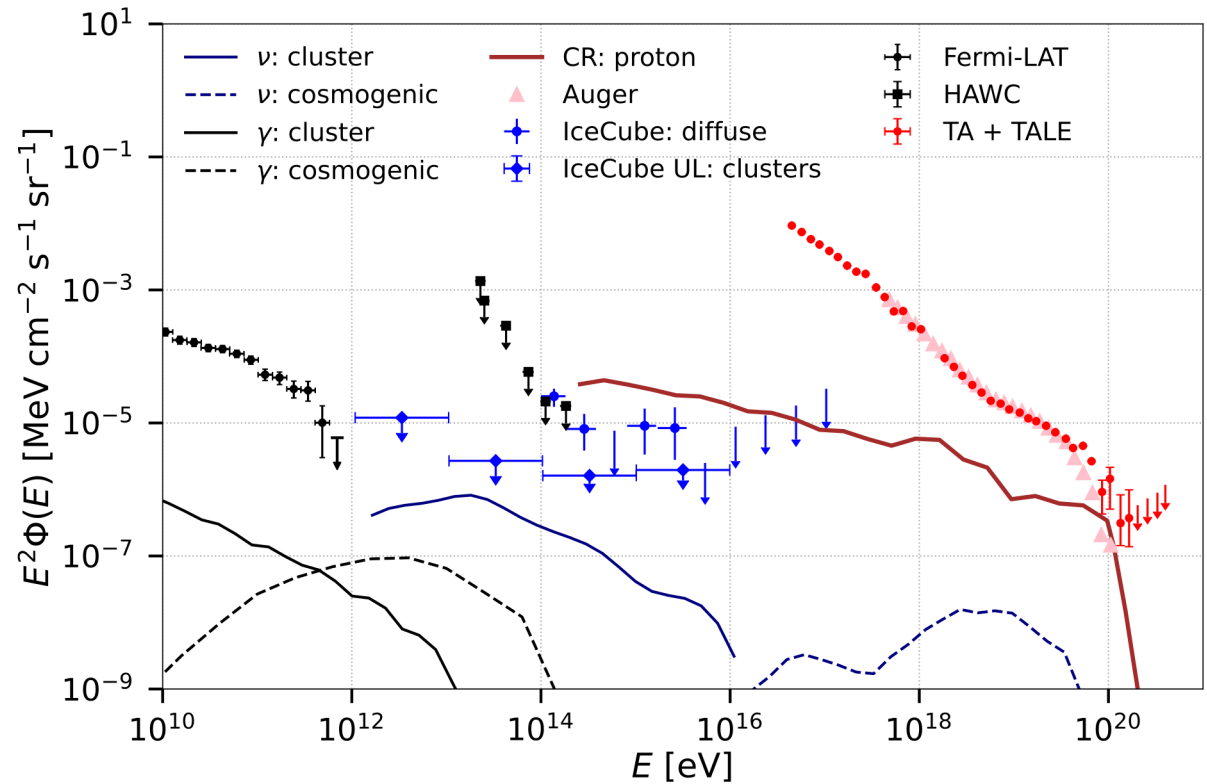


S. Hussain et al. 2026

High Energy Multi-messengers

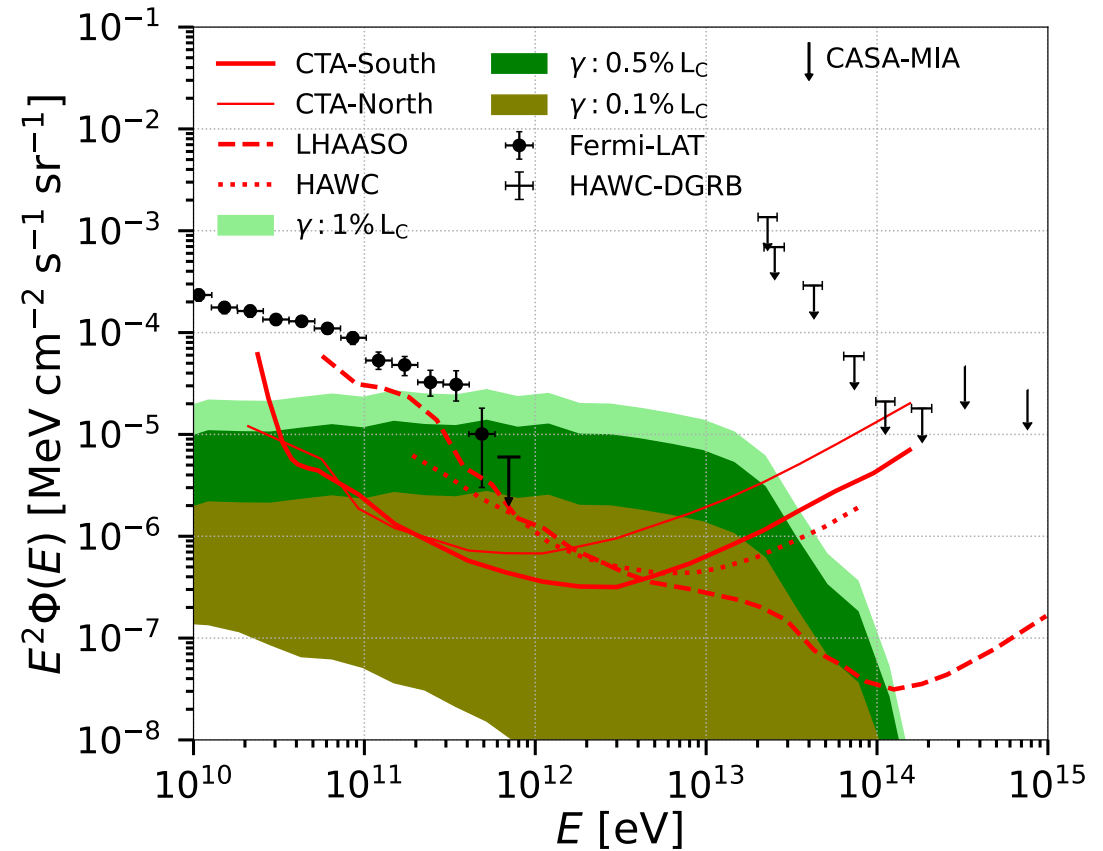
Perseus like source
within 100 Mpc

Spectral index 2.3 &
 $E_{\text{max}} = 1\text{e}17\text{ eV}$



Future Observations

Comparing with sensitivities
Of gamma-ray observatories



Multi-Messenger Testing: Facilities like POEMMA can determine if cluster-center AGNs generate detectable high-energy gamma-ray and neutrino signatures.

Thanks

Currently working

Galaxy cluster

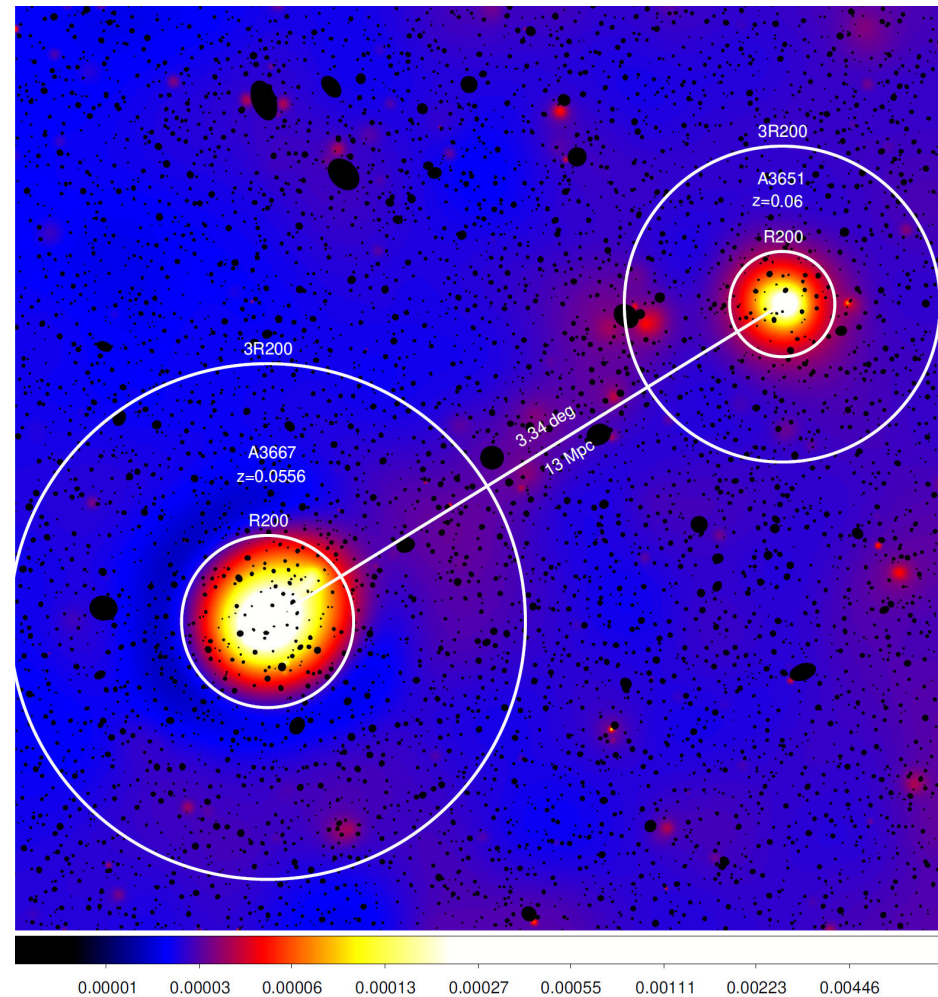
A3667 (SPT-CL J2012-5649) $z=0.0556$

Gamma-ray detection by Fermi-LAT with high significance

Focusing this source

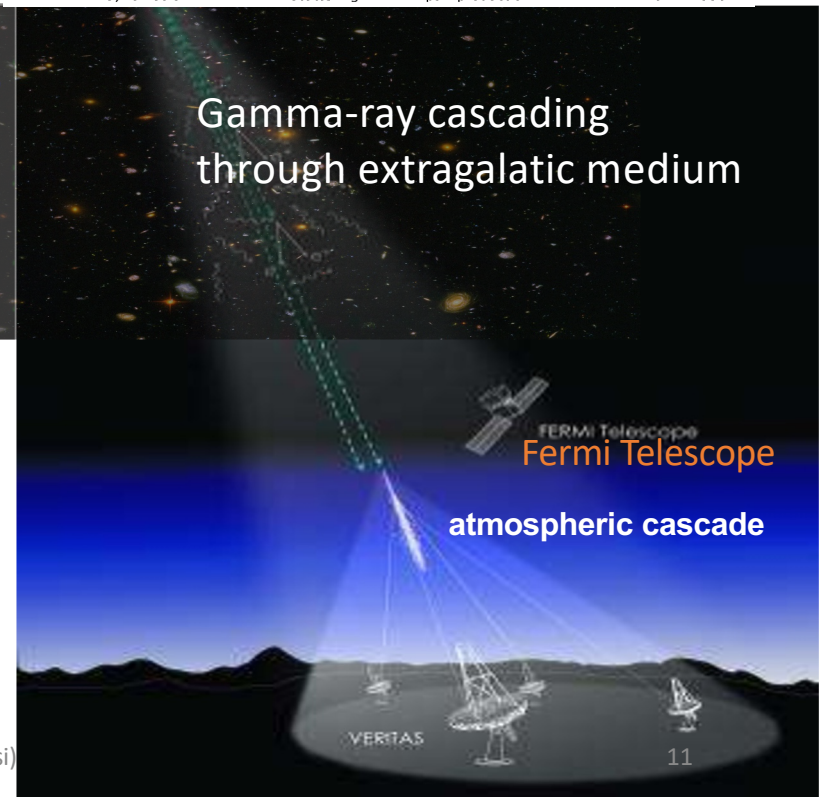
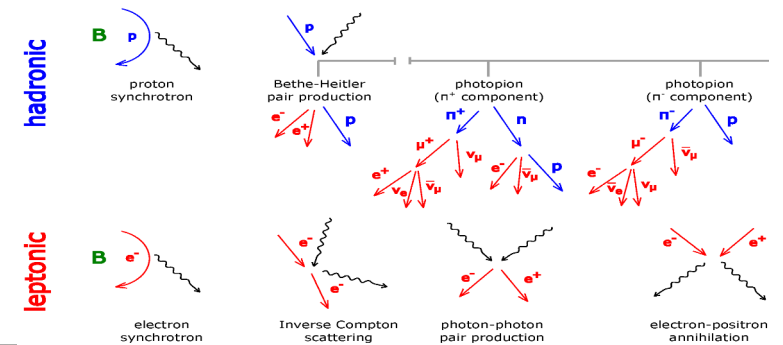
Distinguishing between central source

Outskirt admission.



Extremely long gas filament discovered in X-rays

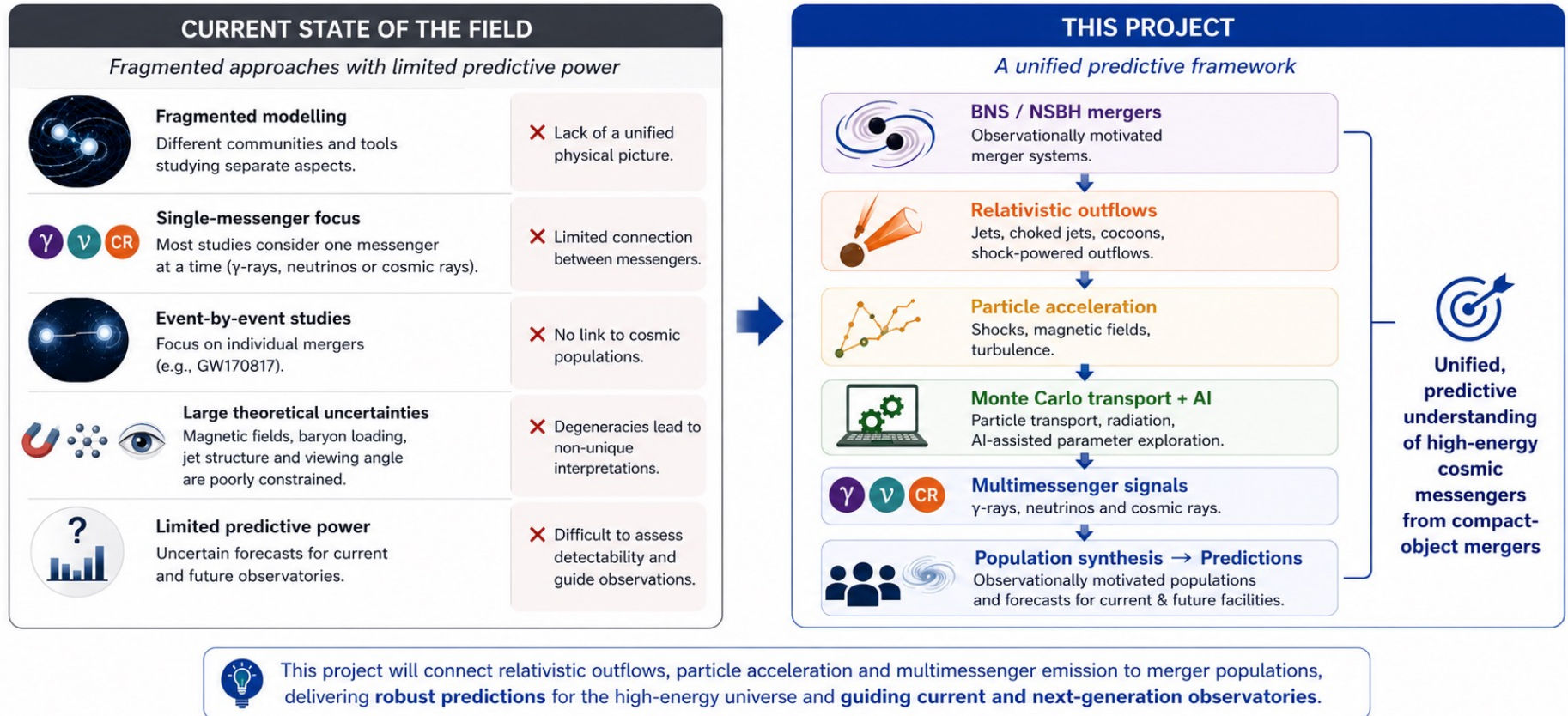
Saqib Hussain (saqib.hussain@ung.si) (Dietl, J., et al., 2024, A&A, 691, A286)



Methodology

Saqib Hussain (saqib.hussain@ung.si)

Currently Working.....



Actively seeking postdoctoral opportunities and collaborative projects in multimessenger astronomy.
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