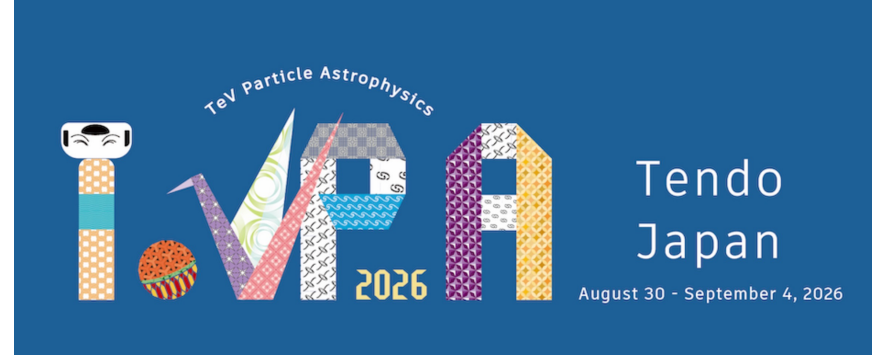
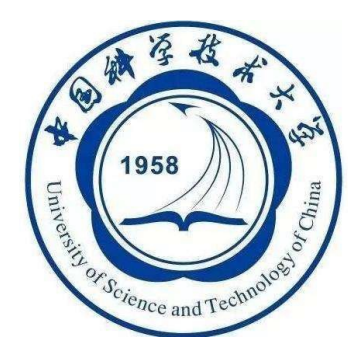




Status and Prospect of the RELICS Experiment

Speaker : Qing Lin (qinglin@ustc.edu.cn)

(On behalf of RELICS collaboration)

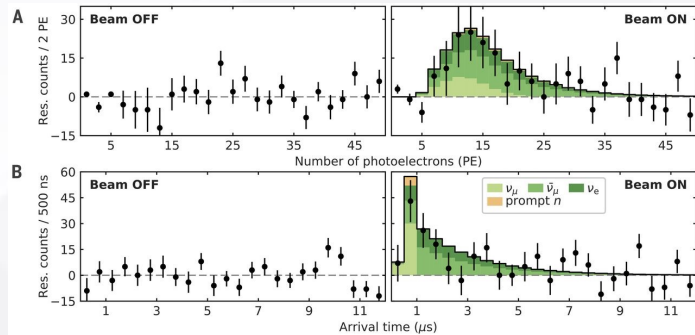
University of Science and Technology of China

TeVPA2026 @ Tendo, 2026.09.02

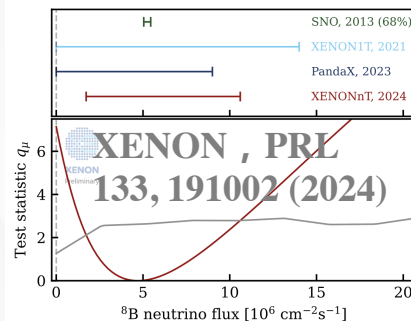
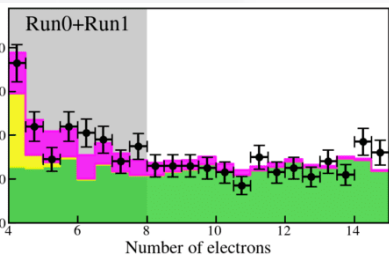
Xe for reactor CE ν NS



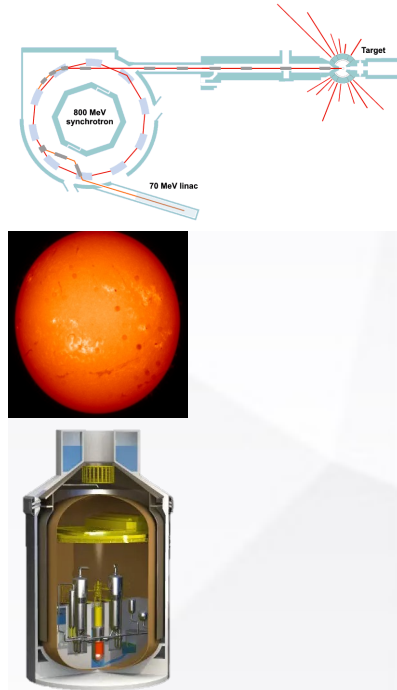
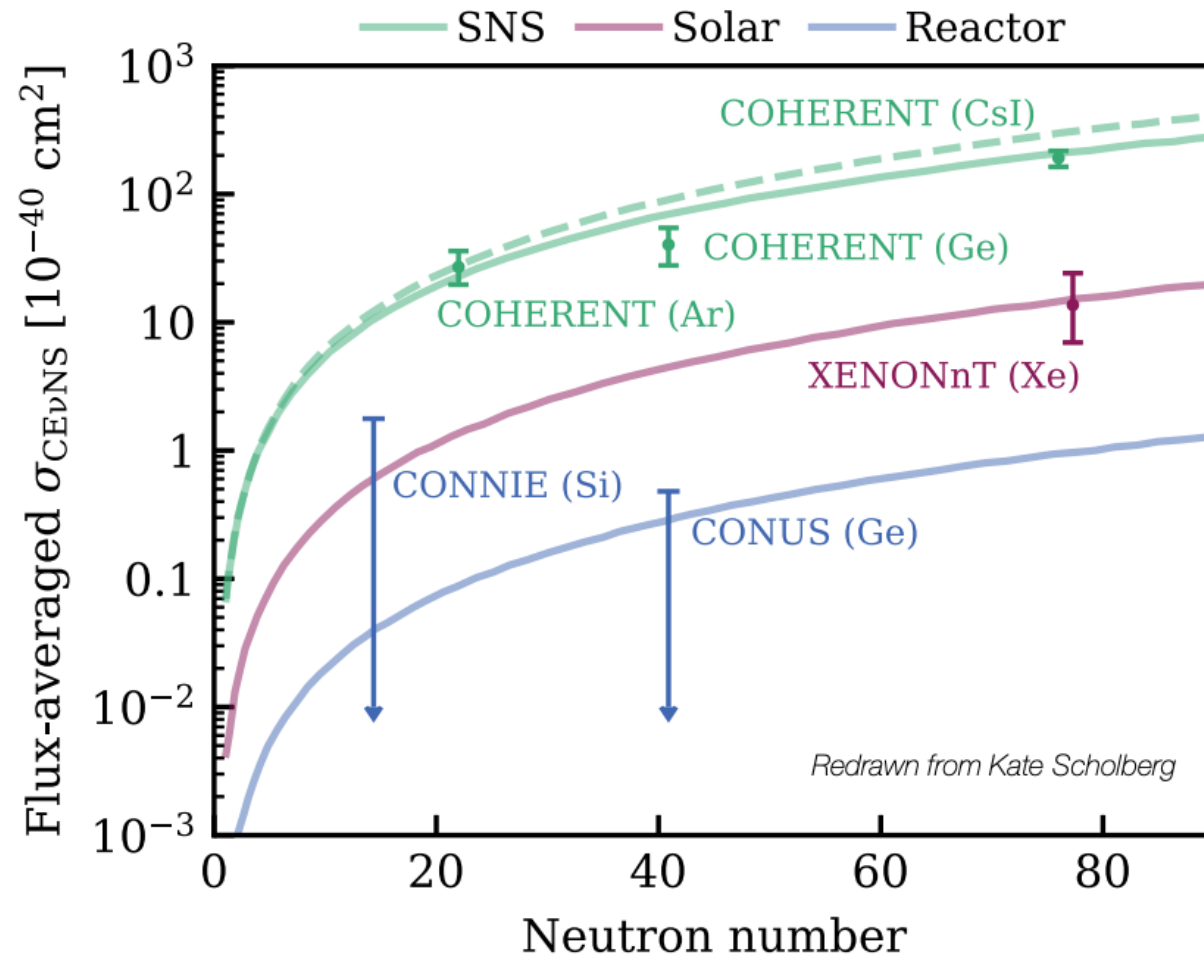
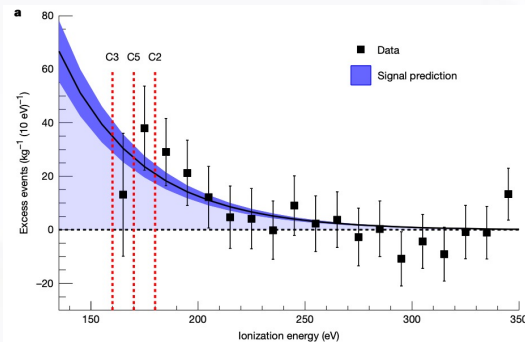
COHERENT, Science 357: 1123 (2017)



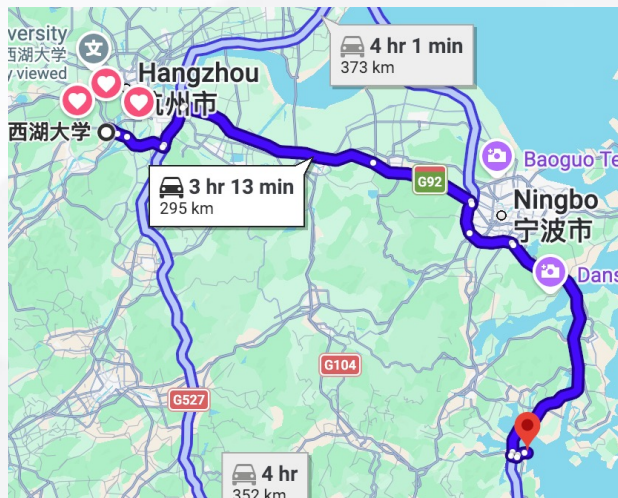
PandaX, PRL 133, 191001 (2024)



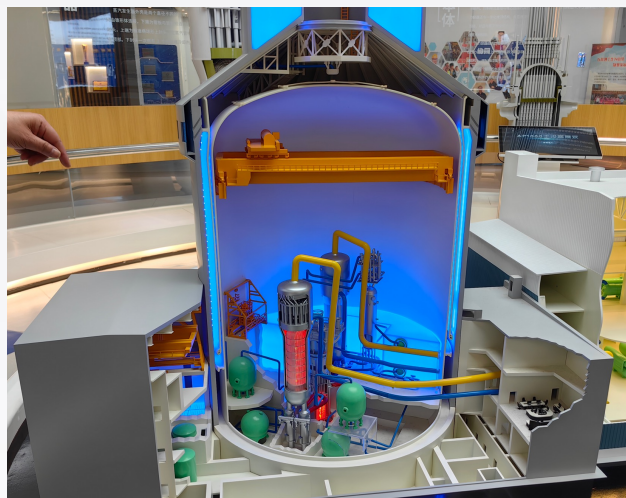
CONUS+, Nature volume 643, pages1229–1233 (2025)



RELICS experiment

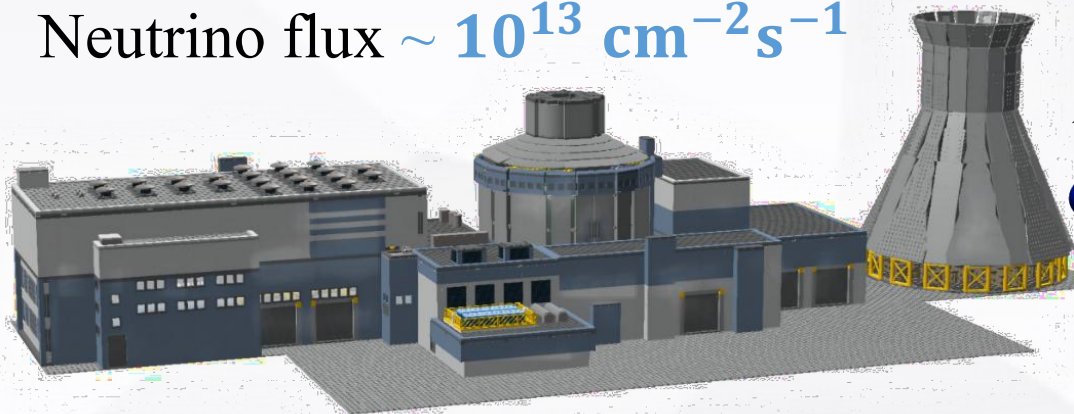


Taizhou, Zhejiang, China

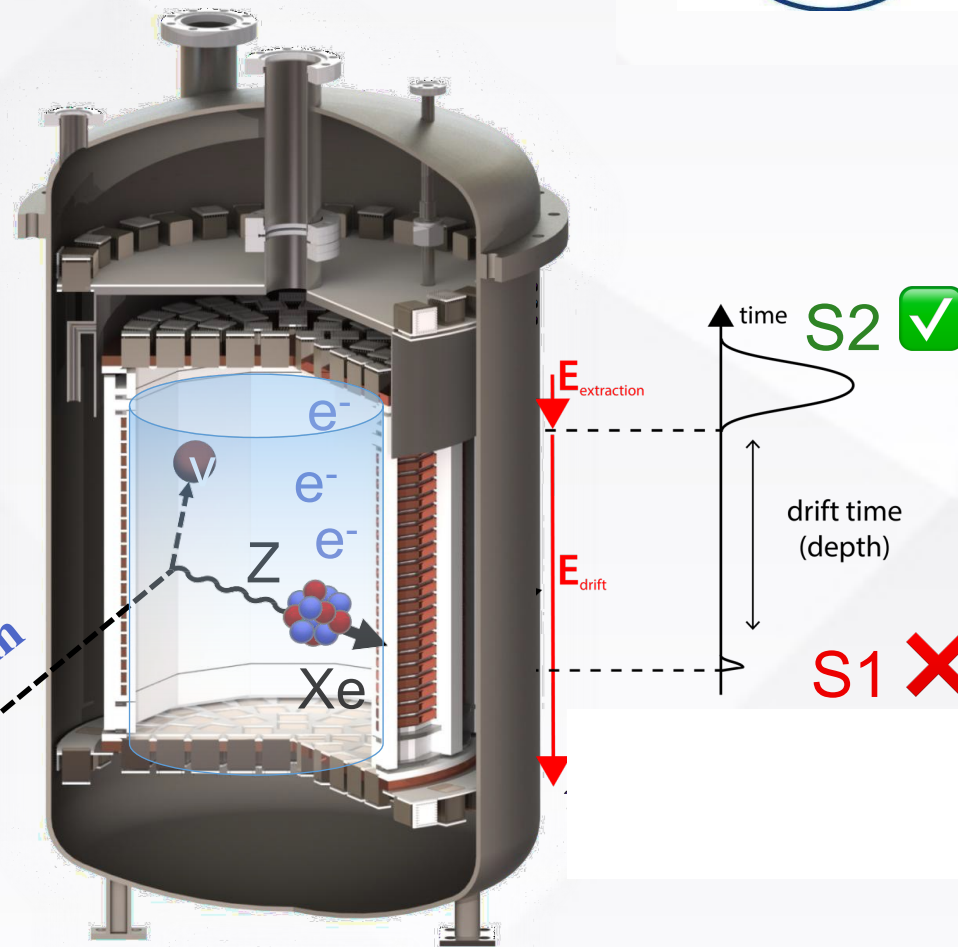


AP1000 thermal power $\sim 3\text{GW}$

Neutrino flux $\sim 10^{13} \text{ cm}^{-2} \text{ s}^{-1}$



Baseline $\sim 25\text{m}$

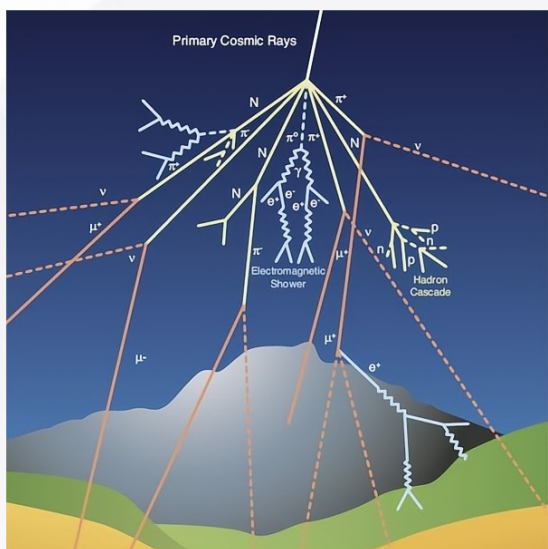


$$13.7\text{eV} = 0.1 \text{ S1} \text{ or } 25 \text{ S2}$$

Reactor neutrino $< 1.3 \text{ keV}_{\text{nr}}$

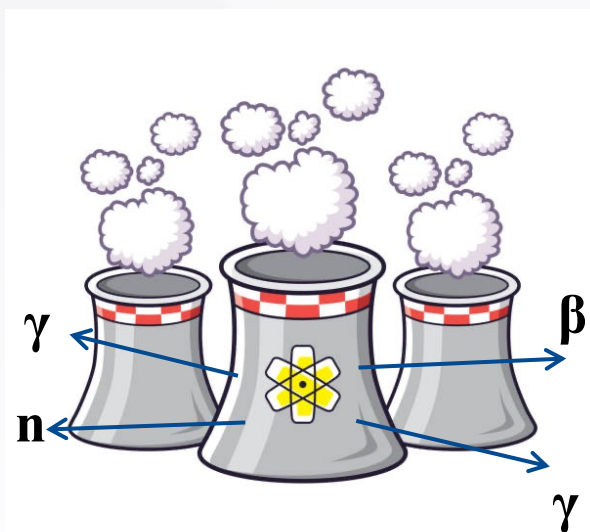
Background Components

Cosmic Ray Particles



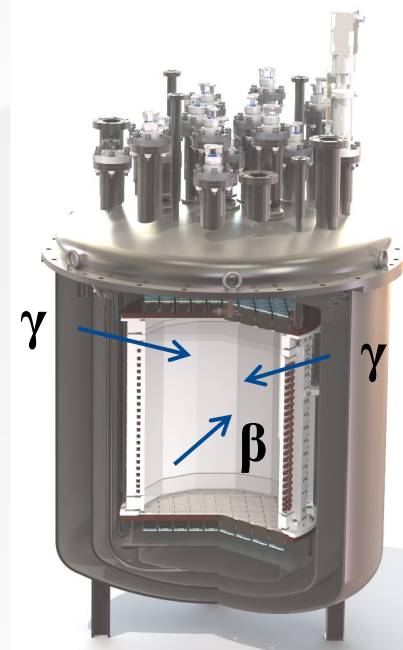
- Cosmic μ
- Cosmic neutron
- Others...

Reactor



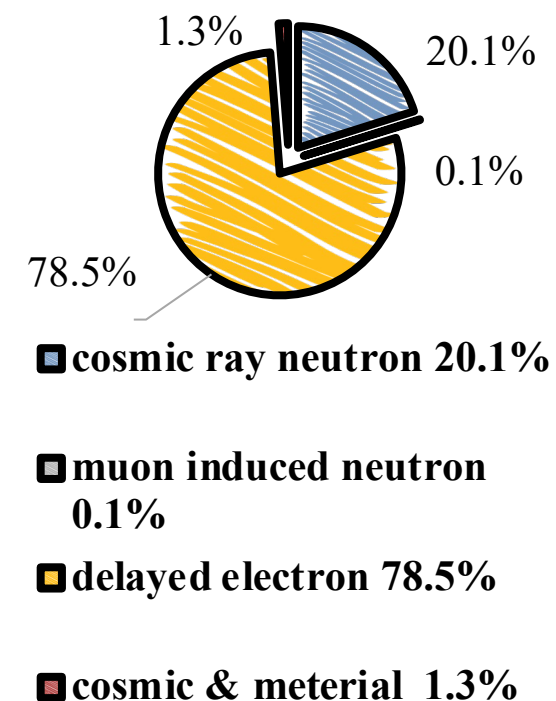
- Reactor Neutron
- Reactor γ
- Others...

Detector Material

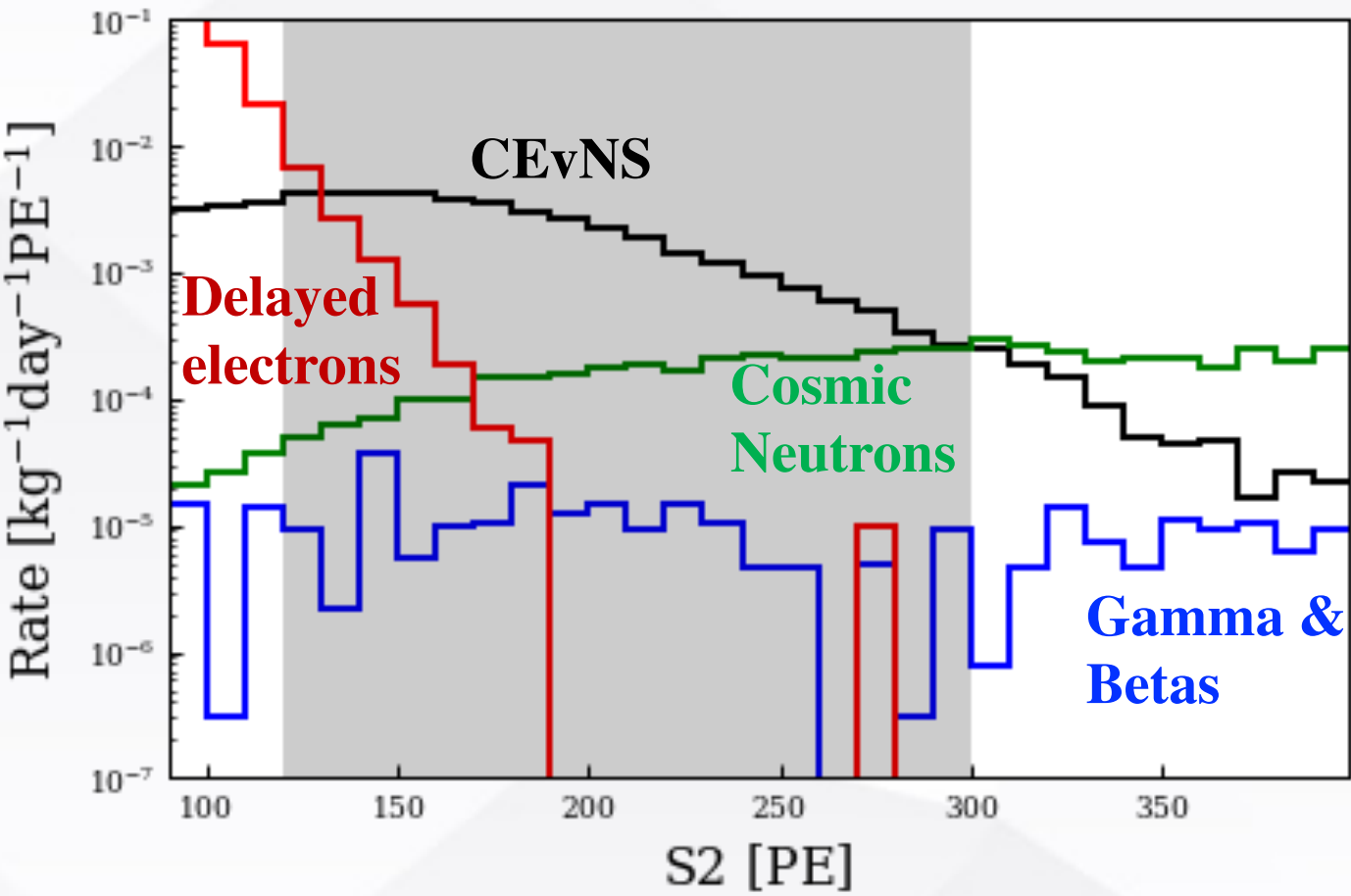


- Material γ , β
- Others...

Background ratio



Signal and Background Expectation



OPEN ACCESS

Reactor neutrino liquid xenon coherent elastic scattering experiment

[Chang Cai¹](#), [Guocai Chen²](#), [Jiangyu Chen³](#), [Rundong Fang⁴](#), [Fei Gao^{1,*}](#), [Xiaoran Guo^{5,6}](#), [Jiheng Guo⁴](#), [Tingyi He^{7,†}](#), [Chengjie Jia^{1,‡}](#)
et al. (RELICS Collaboration)

[Show more](#)

Phys. Rev. D **110**, 072011 – Published 18 October, 2024

DOI: <https://doi.org/10.1103/PhysRevD.110.072011>

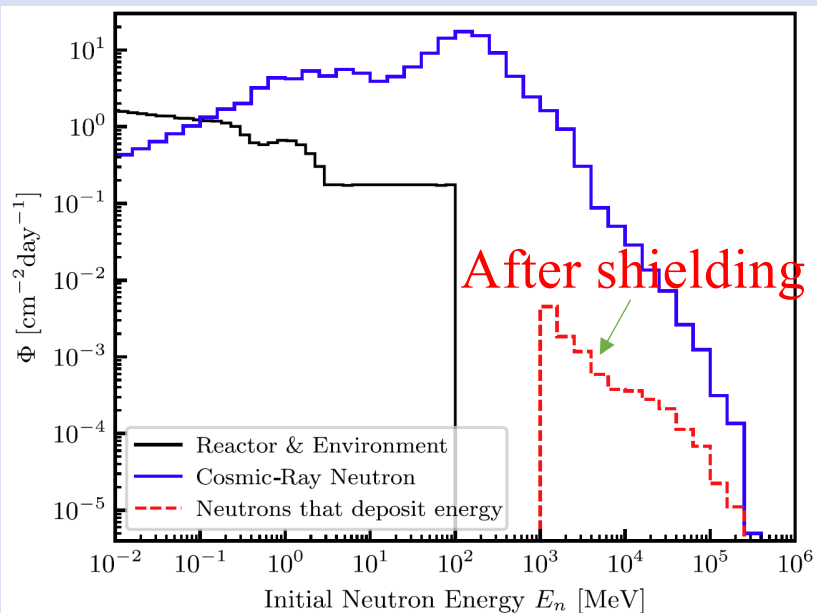
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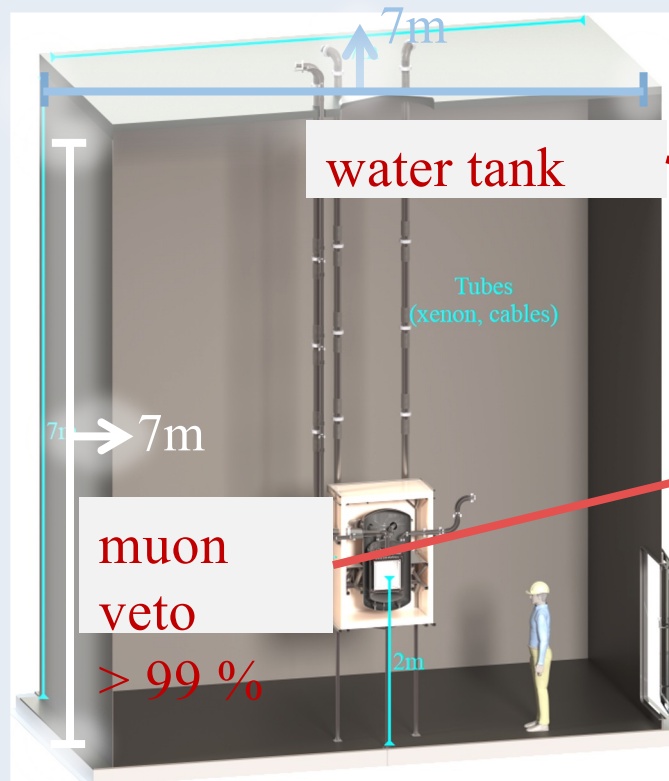
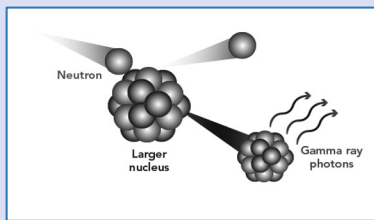
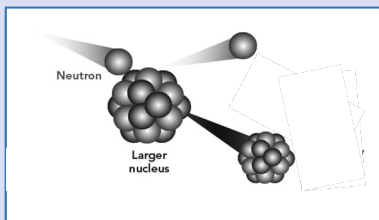
	Events/(32kg · year)
Total background	1687.8
Cosmic neutron	339.9
μ -induced neutron	1.7
Beta, gamma	21.1
Delayed electron	1325.1
CEvNS	4639.7

NR Background

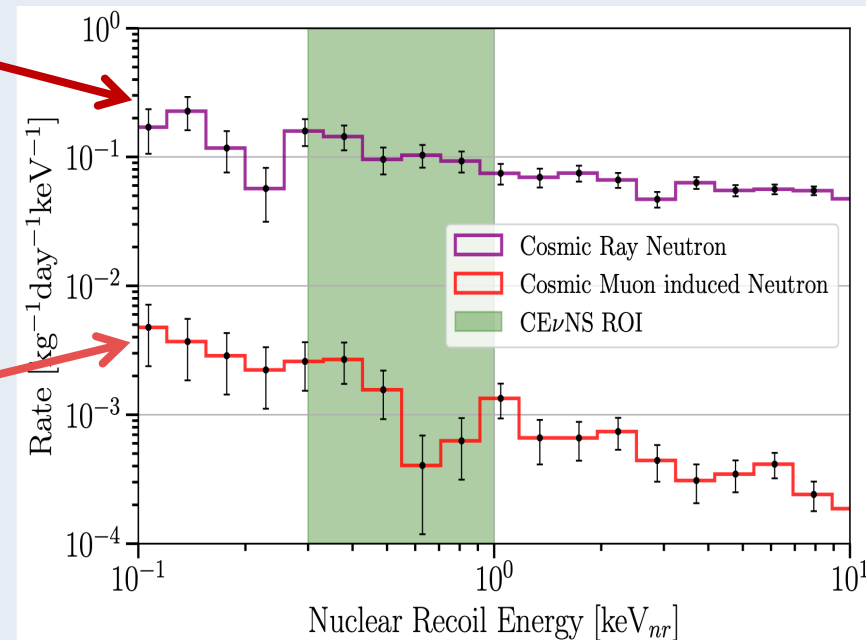


- Cosmic-Ray neutron

- Reactor & Environment neutron



Phys. Rev. D 110, 072011 (2024)

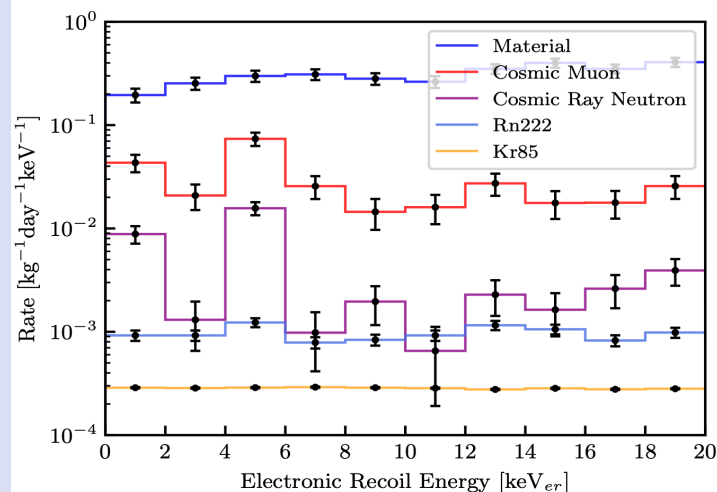


NR Background:

□ $(7.7 \pm 0.7) \times 10^{-2} \text{ kg}^{-1} \cdot \text{day}^{-1}$ in $[0.3, 1] \text{ keV}_{NR}$

□ Dominated by cosmic ray neutron

ER background

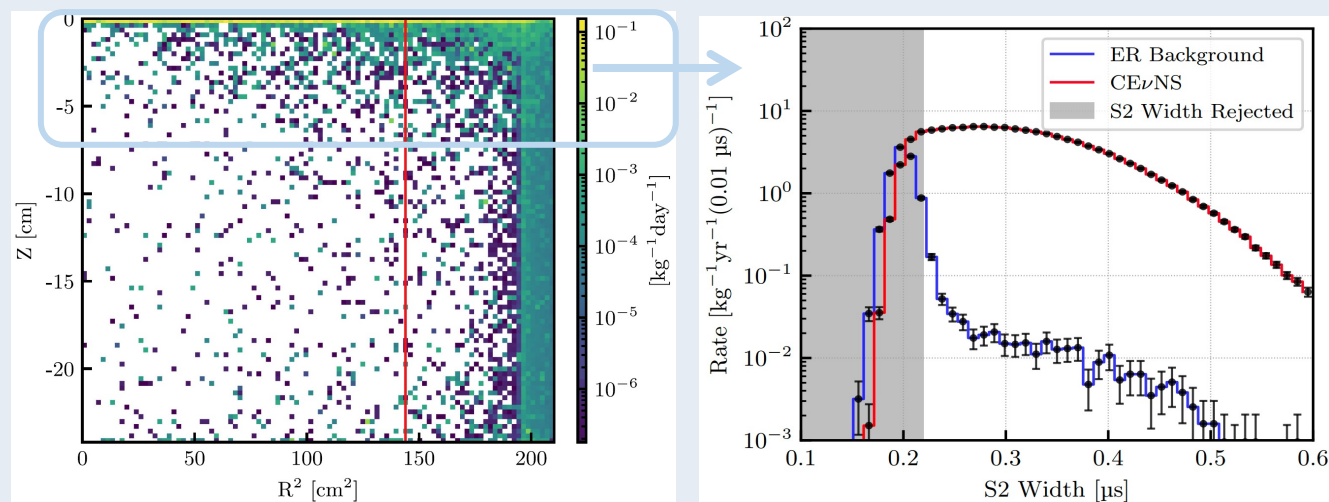
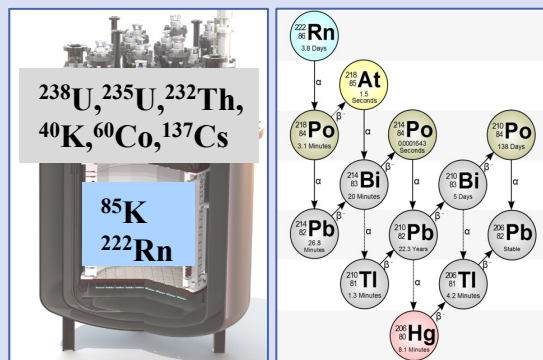
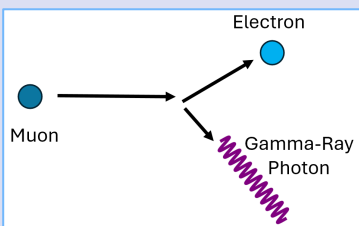


-- Cosmic Ray neutron

-- Material

LXe intrinsic -- Rn222 -- Kr85

-- Cosmic Muon



Fiducial Volume: S2 only -- No Z position

$\boxtimes$ $R < 12\text{ cm}$

$\boxplus$ Edge

> 95% efficiency

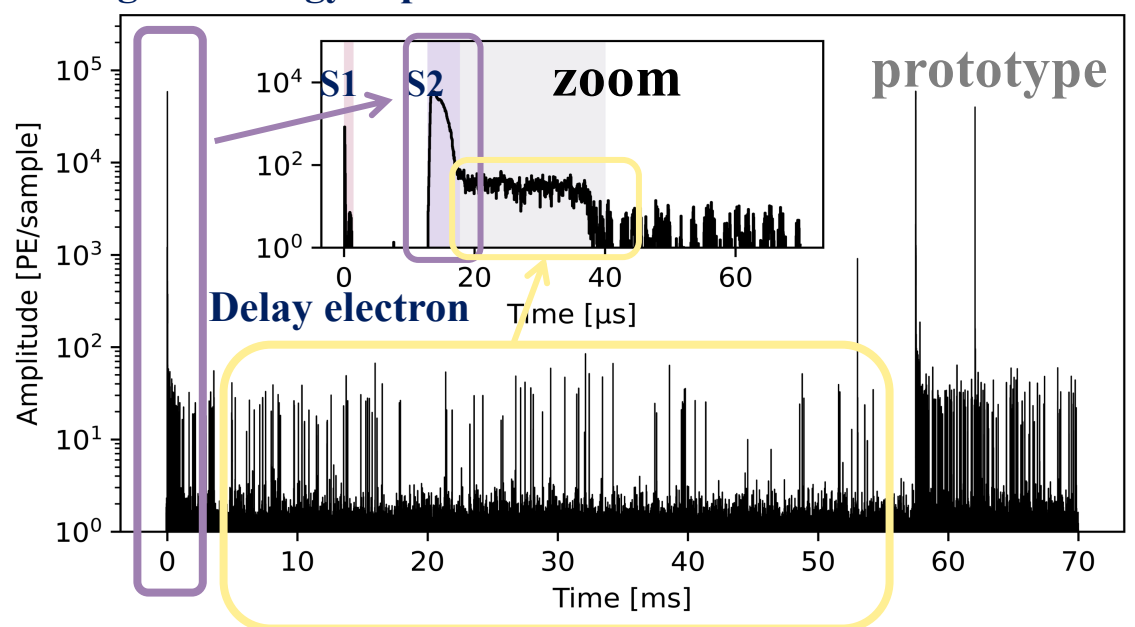
ER Background:

$\square (3.10 \pm 0.10) \times 10^{-1} \text{ kg}^{-1}\cdot\text{day}^{-1}\cdot\text{keV}^{-1}$ average in $[0,20]$ keV_{ER} before S2 width cut

Delayed electron background

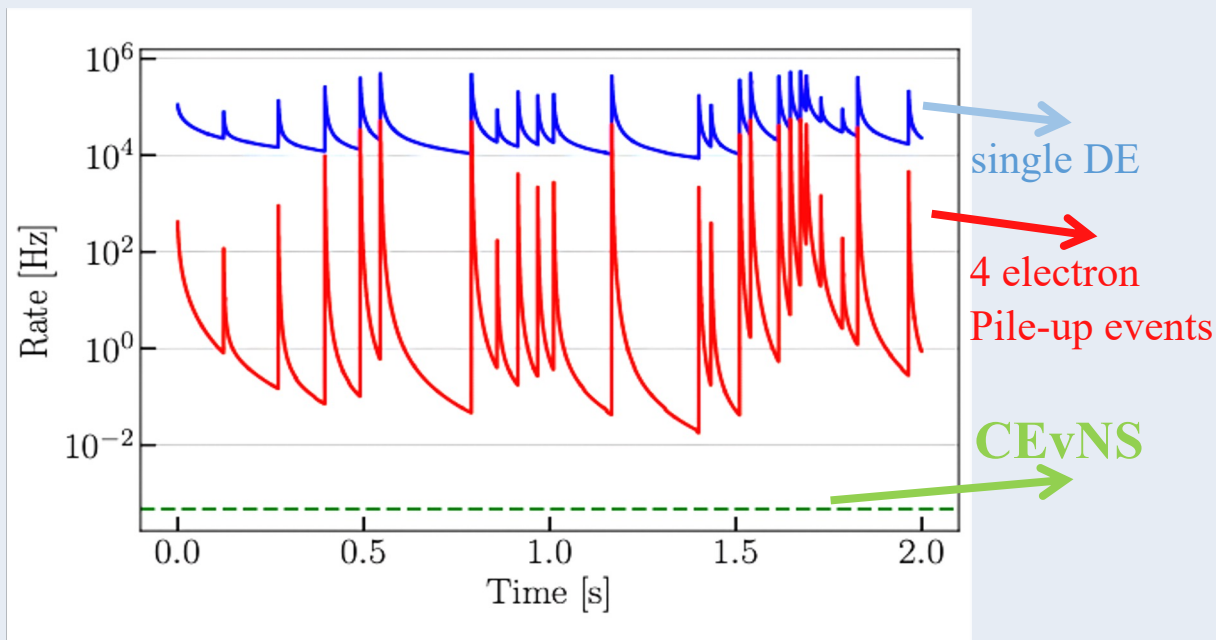


Large S2 energy deposition



Delayed electron sources:

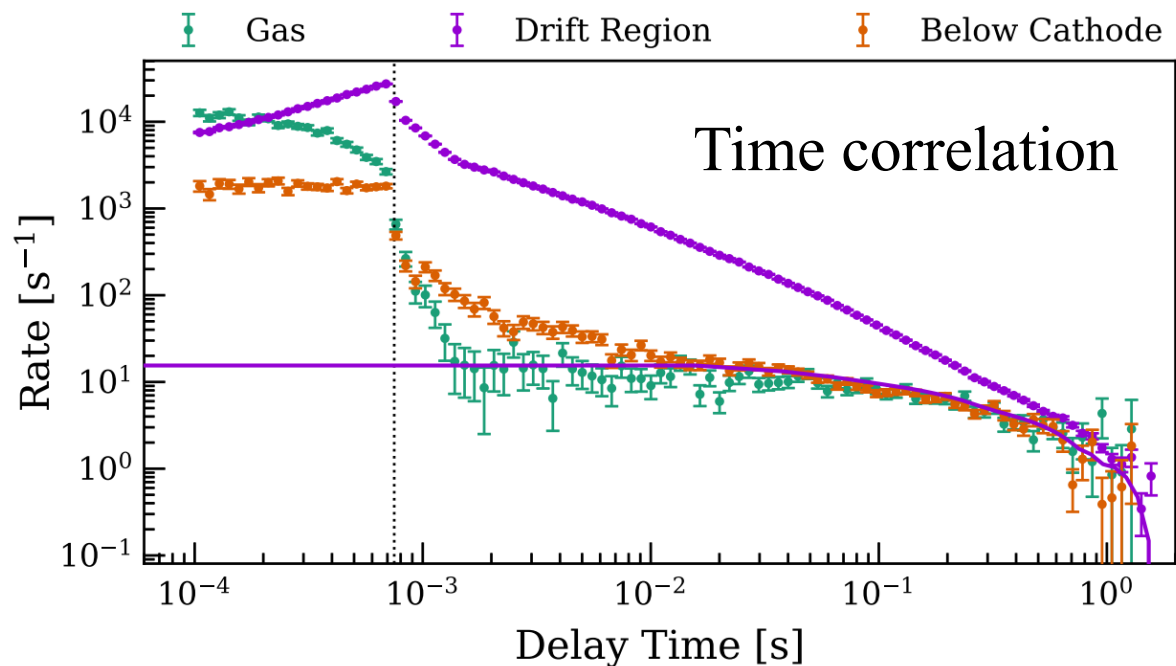
- ❑ Observed by large LXeTPCs experiments and prototypes
- ❑ Emission of electron following large energy deposition
- ❑ Pile-up of single electrons distorts the physical signal



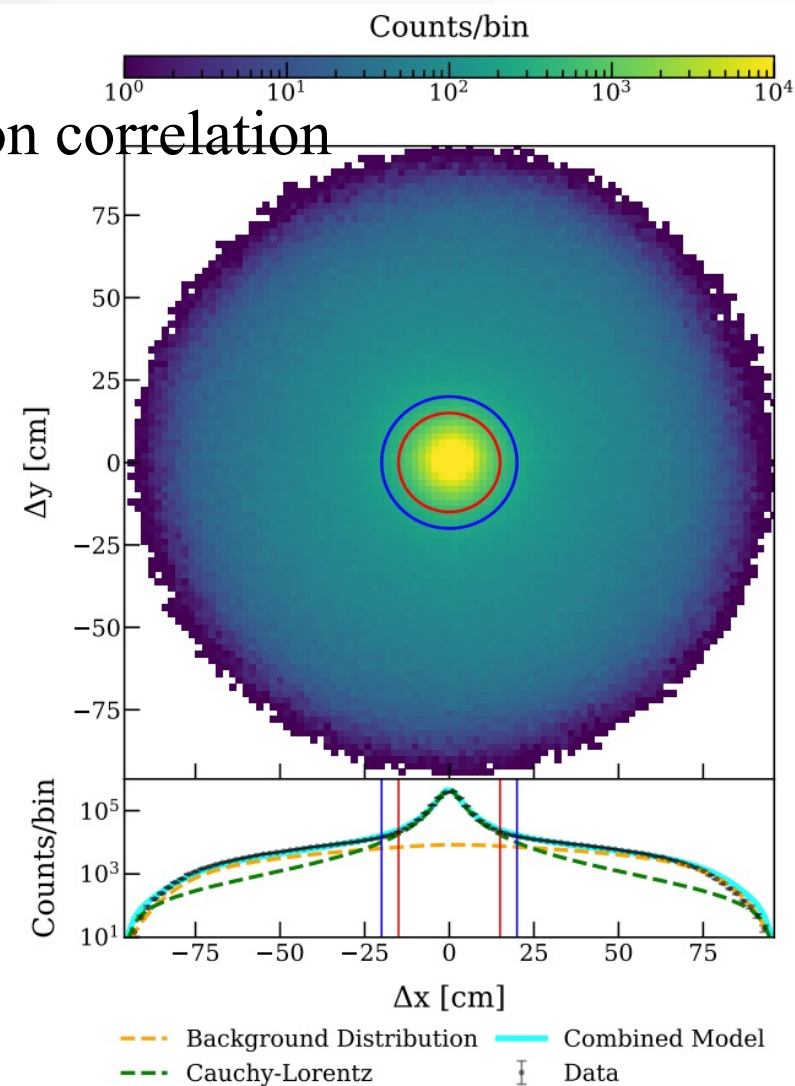
Delayed electron background:

- ❑ ~ 10 Hz muon flux
- ❑ Pile-up DE events rate higher than CEvNS events

Features of DE

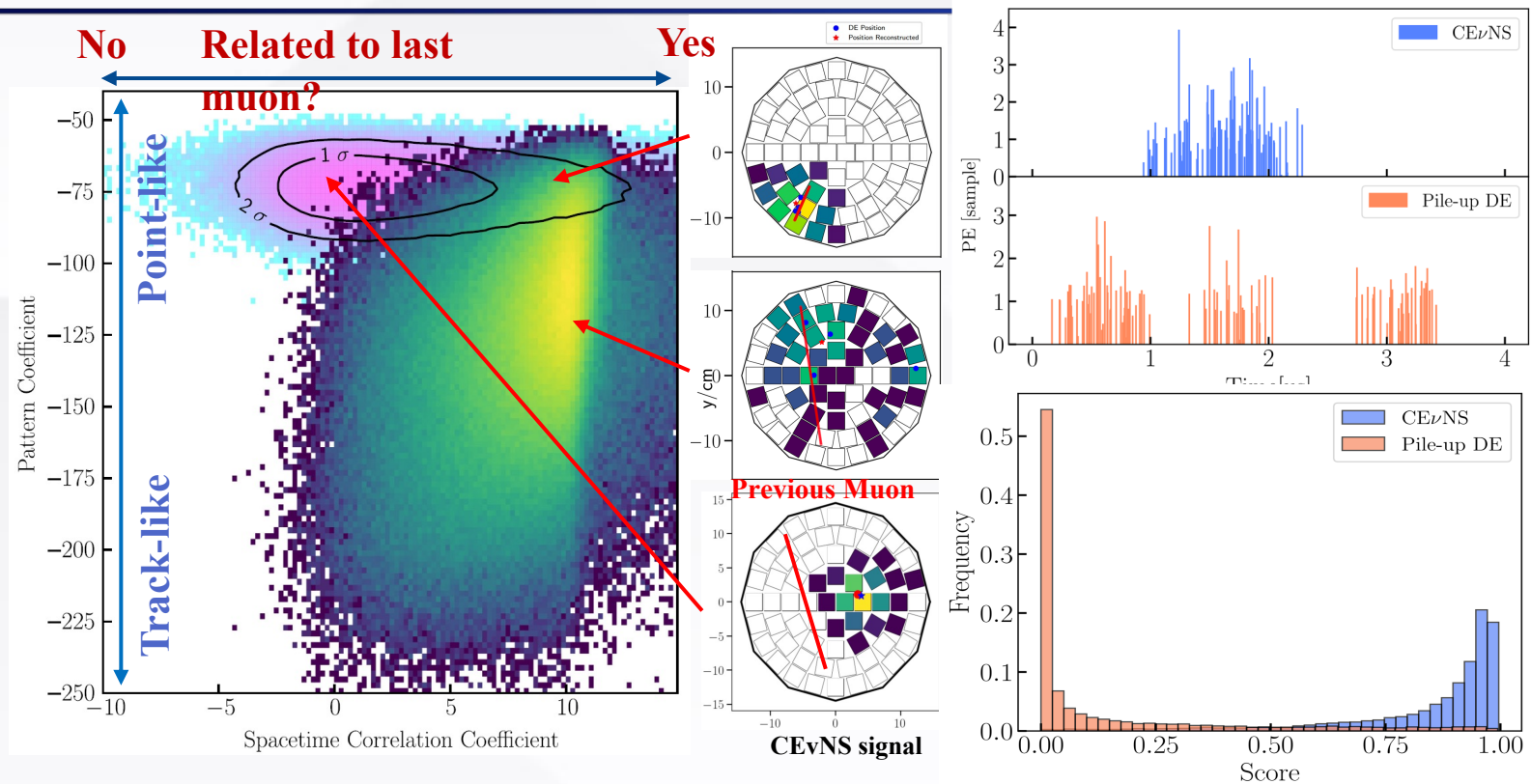


Position correlation



- Delayed electrons studied in large LXeTPCs
- Time correlation
- Position correlation

DE Suppression



	Signal Acceptance	Background Remaining
Pattern + Correlation	~52%	~0.01%
Waveform	~80%	~10%

4-5 orders total

Pattern + Correlation selection:

- CEvNS : Closer to single - point events.
- DE: More linked to the preceding muon track.

Waveform selection:

- CEvNS : clustered
- DE: dispersed

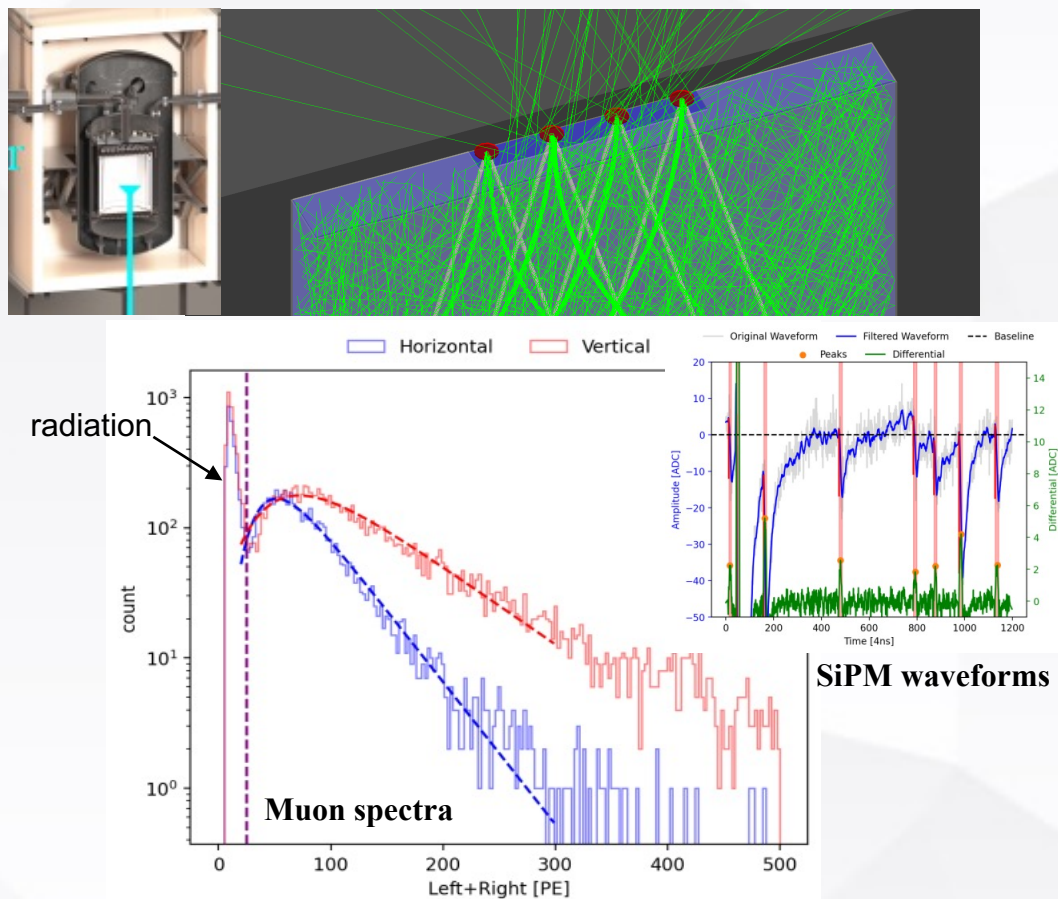
Need:

- Full muon track information
- DE suppression with material

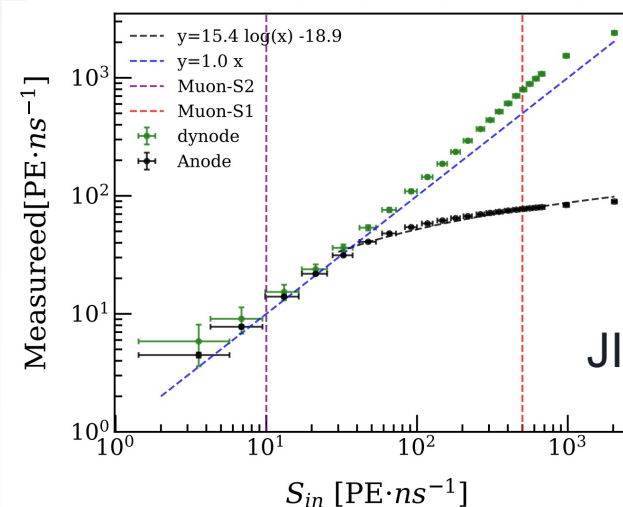
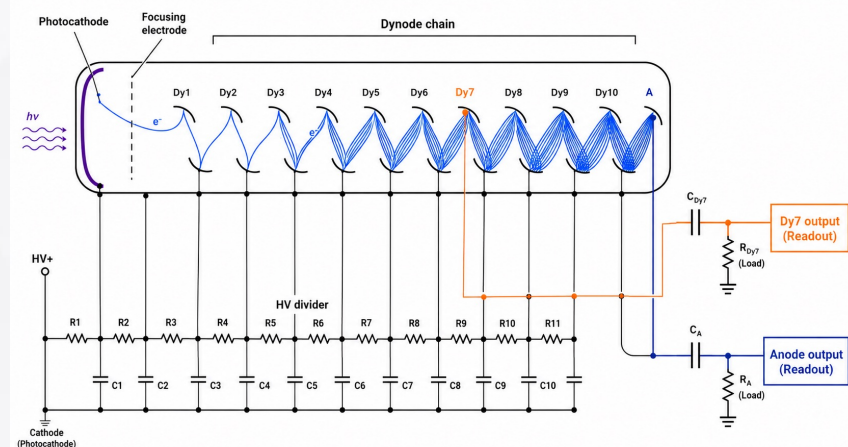
Cosmic muon detection



Path 1: Muon **Active Veto** Boards



Path 2: PMT **Dynode** Readout Extension



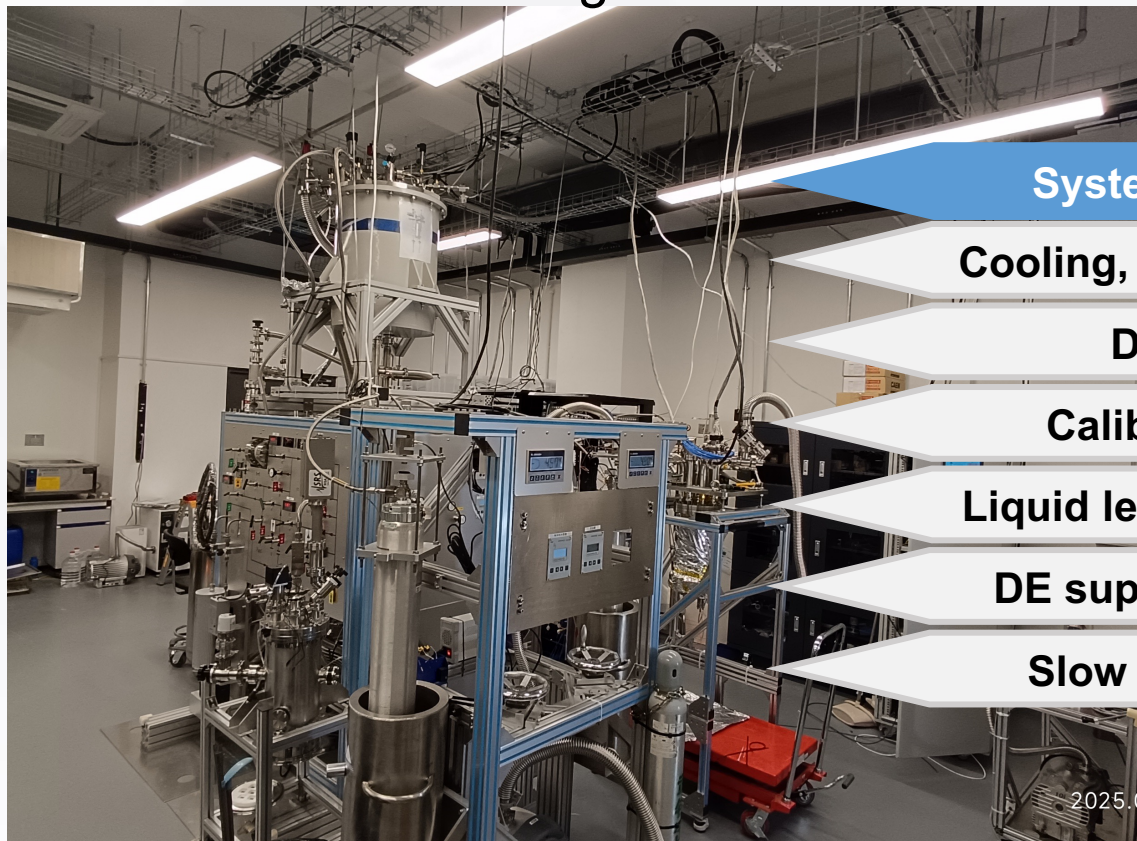
JINST 21, P02039 (2026)

Range: keV $\rightarrow$ MeV
Dynode recovers otherwise **saturated** anode signal

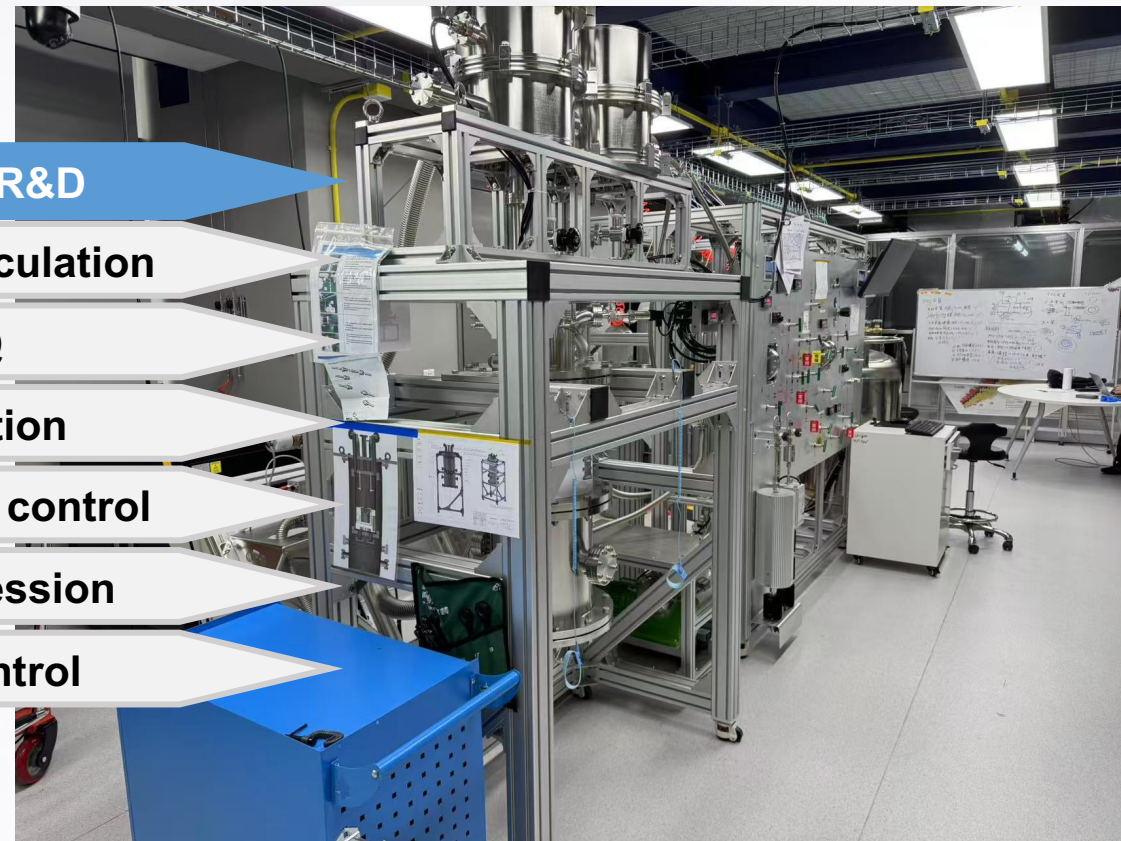
RELICS prototype



Tsinghua



Westlake



System R&D

Cooling, circulation

DAQ

Calibration

Liquid level control

DE suppression

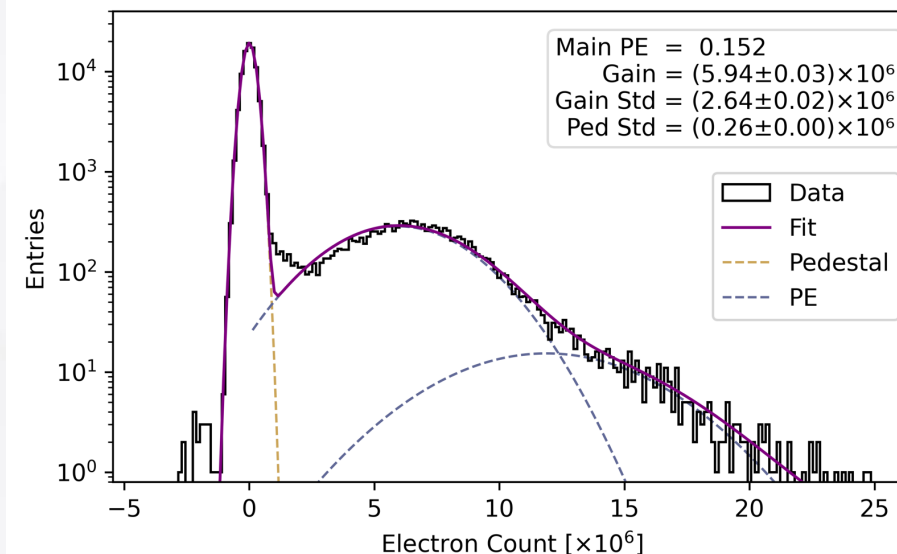
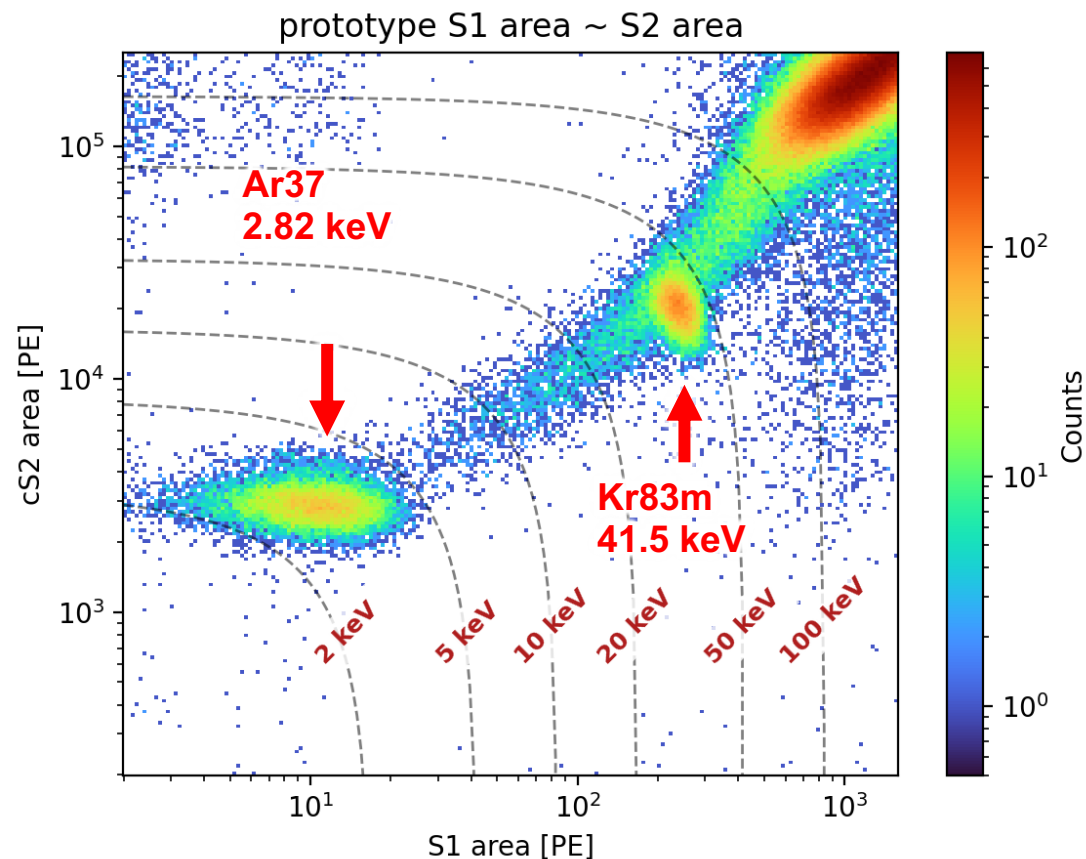
Slow control

NST 37(1), 14 (2025)
Eur. Phys. J. C 86, 348 (2026)
arXiv:2604.12572

Test and optimize sub-systems
Establish data analysis pipeline

JINST 21, P02039 (2026)

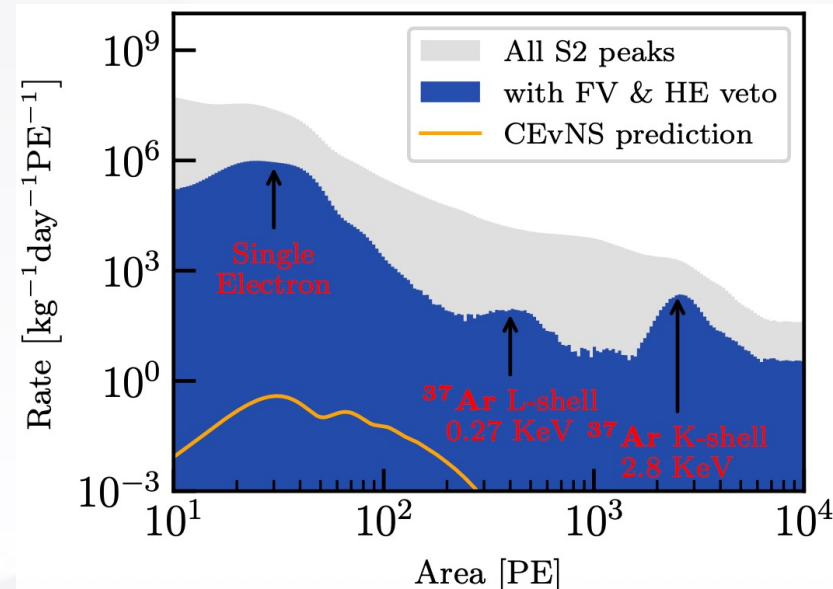
Energy threshold and DE suppression



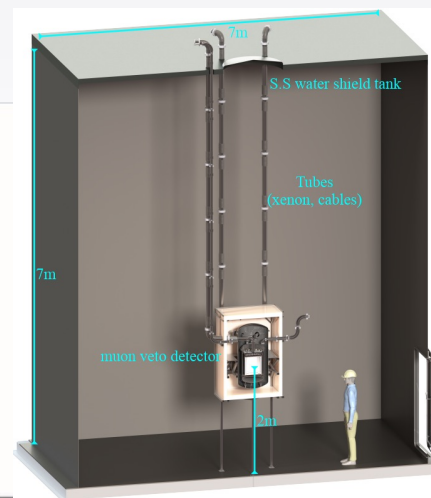
Single electrons are detected with high efficiency
DE background can be suppressed via HE veto

NST 37(1), 14 (2025)

Eur. Phys. J. C 86, 348 (2026)



First step: RELICS-10

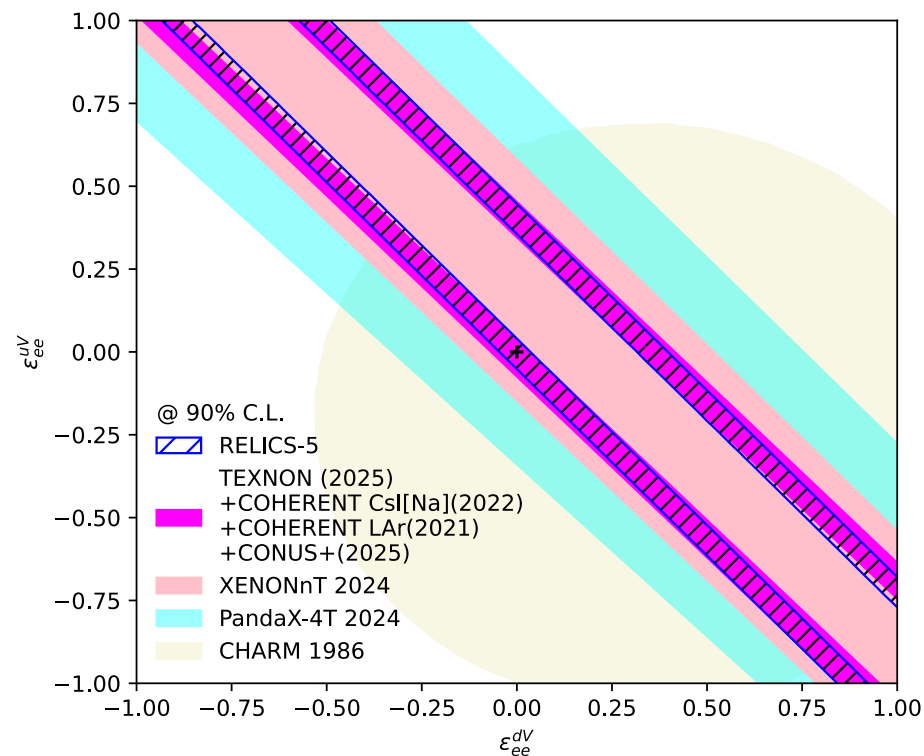
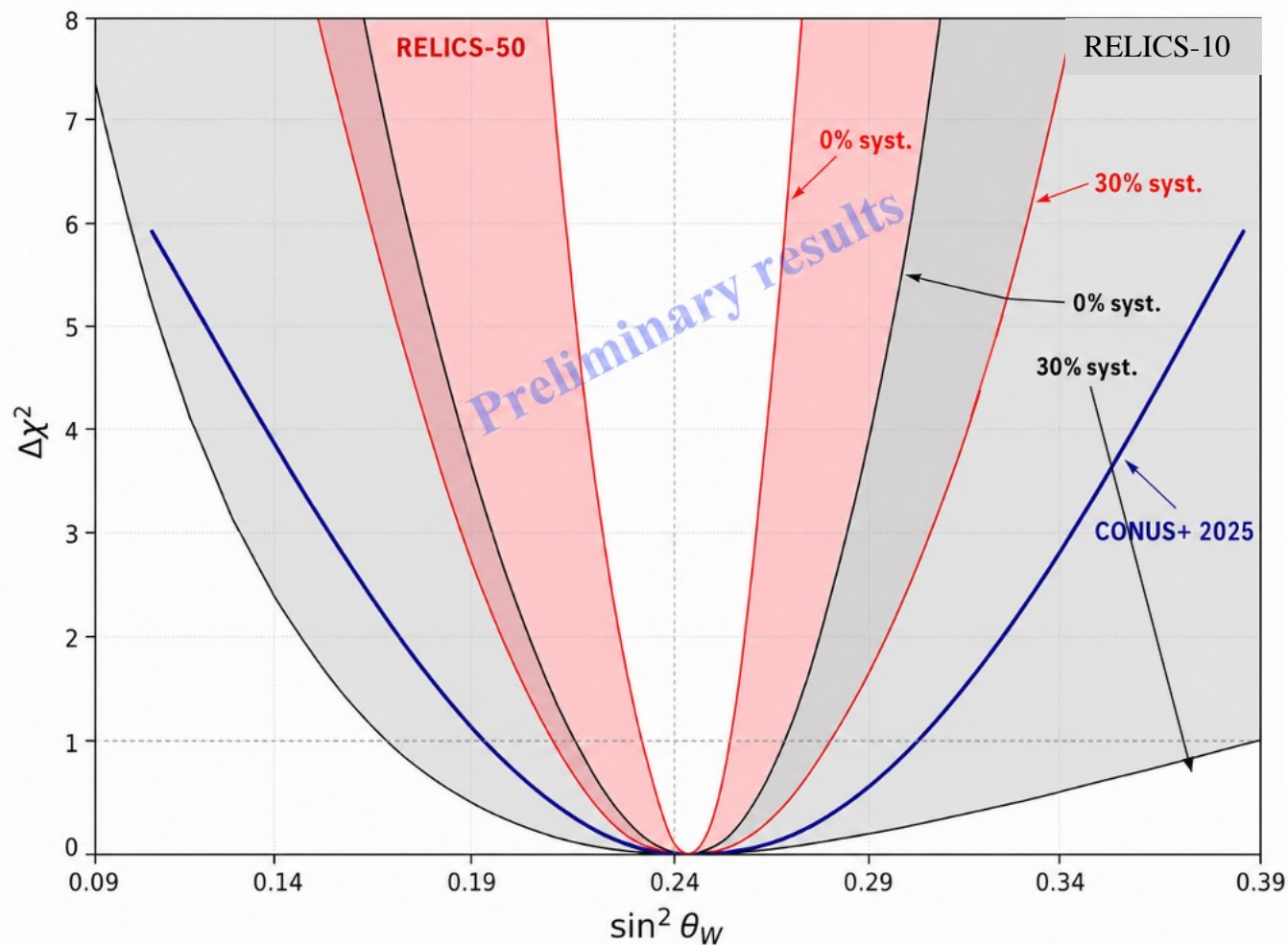


Increased background ➡ smaller detector

	Events / (5 kg·year)
CEvNS	4640 x 0.15 ~ 720
Cosmic Ray Neutron	340 x 3 ~ 1000
Muon Induced Neutron	2 x 1 ~ 2
ER	21 x 5 ~ 100
DE Pile-ups	1325 x 0.33 ~ 440

Sanmen 2027

Physics potential for RELICS-10



RELICS-10 is competitive in CEvNS physics

Timeline



Tsinghua



Westlake



Reactor 



PROTOTYPE

RELICS-10

RELICS-50

2024

2025

2026

2027

2028

2029

Subsystem R&D

Commissioning

Next step

Summary



- RELICS aims to measure reactor CEvNS using a **liquid-xenon TPC**.
- RELICS **prototypes** and **subsystems** were built to demonstrate the **feasibility** of the detector concept.
- **RELICS-10** will commission in **2026**, and target CEvNS detection in **2027**.



Backup

